

# X-RAY LASER PHYSICS

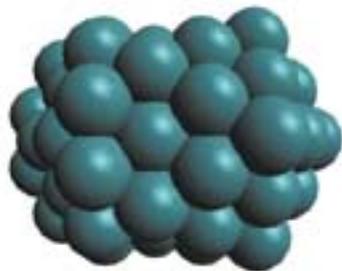


Shanghai Jiao Tong University, Shanghai, China

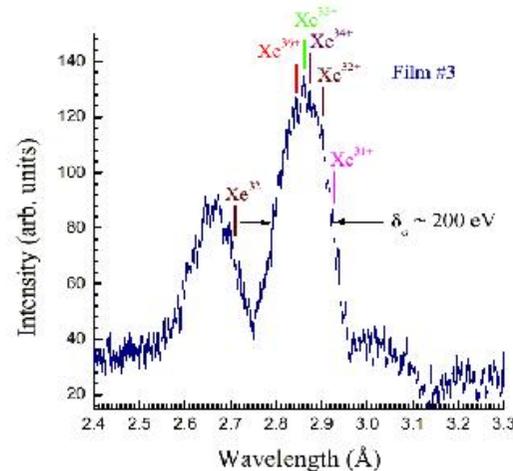
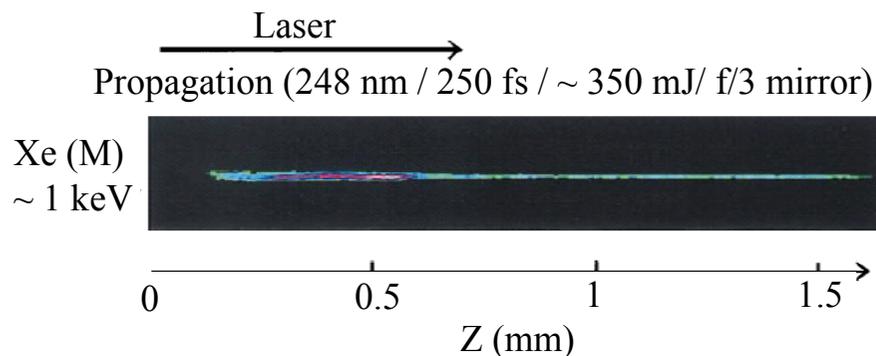
Charles K. Rhodes  
Jiao Tong University  
Shanghai, China  
October 27-31 2008

# X-Ray Amplification

## Two Key Innovations



- **Channeled Propagation**  
→ **Spatial Organization**



## Historical Paper

In their seminal paper “Infrared and Optical Masers” published in *The Physical Review* in 1958, Arthur L. Schawlow and Charles H. Townes concluded with a section in which they discussed the high-frequency limits of amplification [1]. They wrote “unless some radically new approach is found,” the x-ray zone is out of reach. Furthermore, the last sentence of the same article stated “...continuous tuning over larger ranges of frequency will require materials with very special properties.” In relation to the work described below, the former requirement is satisfied by the development of a new concept that involves radiation dominated energy flow under conditions for which the phase space of the interaction is precisely controlled while the latter is met by the use of hollow atoms.

[1] Townes and Schawlow, Phys. Rev. **112**, 1940 (1958)

**Key Concept → Controlled Power Compression**

# Why X-Rays?

## Power Density Scaling

### → TRANSITION TO X-RAY REGIME

- Basic Parameter (**wavelength**) Changed by  $> 10^3$

$$1 \text{ eV (IR)} \xrightarrow[\sim 4500]{\lambda^{-1}} 4.5 \text{ keV (X-Ray)}$$

- Radiative Rates Scale Like  $\sim \lambda^{-3}$

$$(4500)^3 \sim 9 \times 10^{10}$$

- Area Scales  $\sim \lambda^{-2}$

(Channeling)

$$(4500)^2 \sim 2 \times 10^7$$

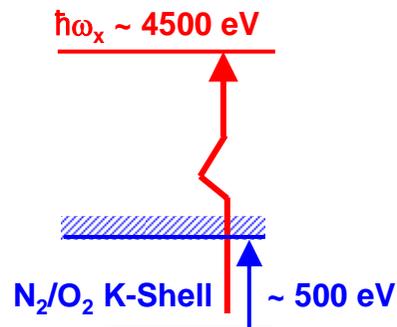
- Combined Power Density Scaling

$$\xrightarrow{\lambda^{-6}} (4500)^6 \sim 8 \times 10^{21}$$

- **HUGE FAVORABLE POWER COMPRESSION SCALING FACTOR ( $> 10^{21}$ )**



### ★ NEW SITUATION WITH LOW-Z MATERIALS



- High Energy  $4.5 \text{ keV} > 300 \text{ Ry} > \alpha m_e c^2$
- High Speed  $\tau_x \sim < 10^{-17} \text{ s} < \hbar/\alpha^2 m_e c^2 \sim 24 \text{ as}$
- High Intensity  $I_x \sim 10^{26} \text{ W/cm}^2$ ,  $E_x \sim 10^5 \text{ e/a}_0^2$
- $r_x \leq \lambda_x \sim 5 a_0$  (molecular scale)

- ★ Enables the creation of new compact ordered states of matter
- ★ Control  $e^-$  Motions

# EXPERIMENTAL CONFIGURATION

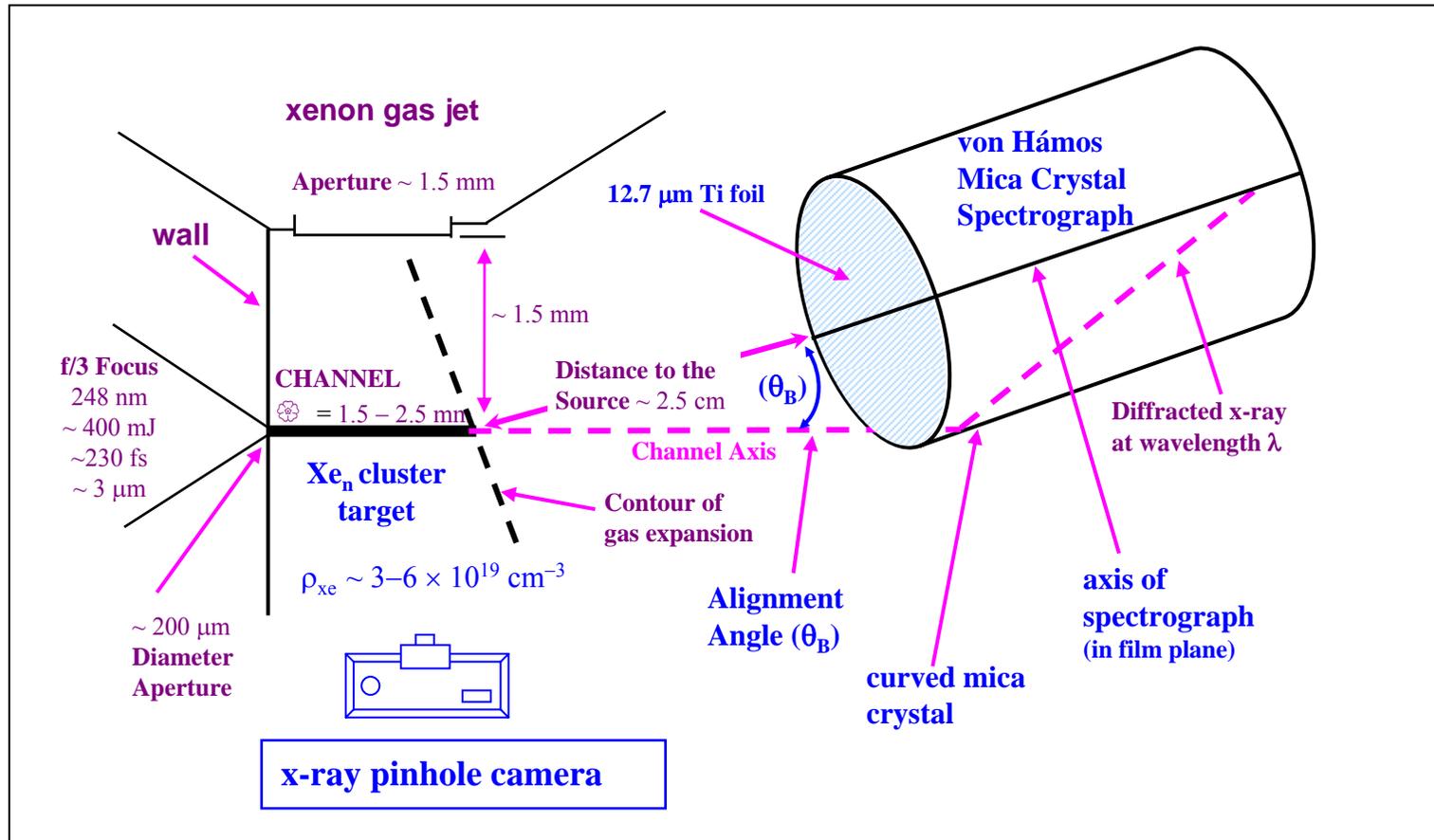
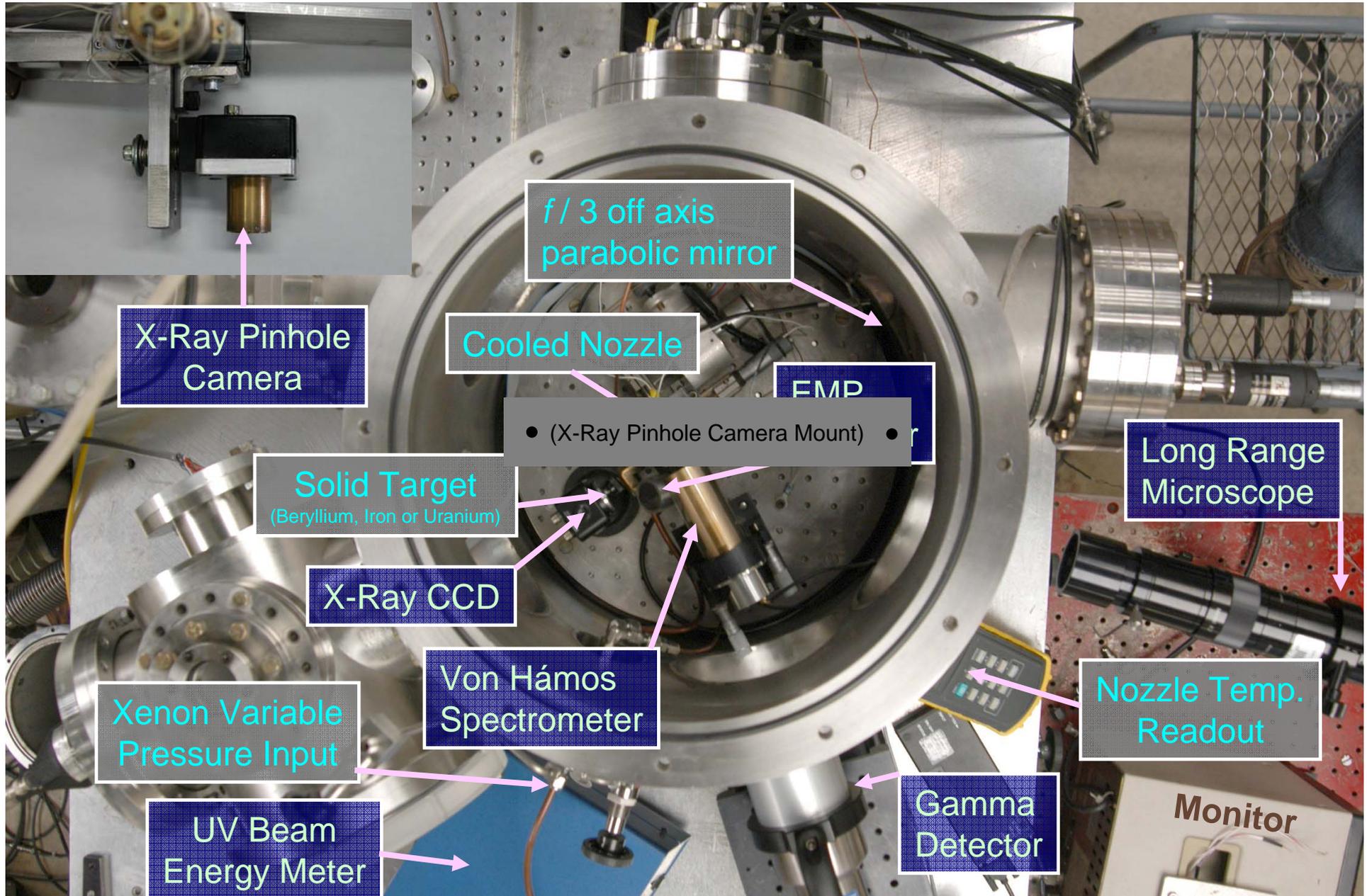


Fig. (2): Experimental configuration used for the observation of amplification of Xe(L) radiation in self-trapped channels inside an evacuated chamber. The x-ray pinhole camera was equipped with a ~ 10 μm thick Be foil enabling the morphology of the channel to be visualized by the Xe(M) emission (~ 1 keV). The observed channel length typically is  $\varnothing \cong 1.5\text{--}2.5$  mm. The wall defining the entrance plane having the 200 μm aperture was fabricated from ~ 100 μm thick steel and the incident 248 nm pulse was focused with an f/3 off-axis parabolic optic to a spot size of ~ 3 μm. The entrance of the von Hámos spectrograph viewing the forward directed emission was protected with a Ti foil of 12.7 μm thickness whose transmission factor in the 2.7–3.0 Å region is ~ 0.5. The Bragg angle for the Xe<sup>34+</sup> component at 2.88 Å is  $\theta_B \cong 26^\circ$ . The film plane, which lies on the axis of the instrument, does not have a direct path to the x-ray source and, hence, only receives exposure by diffraction from the curved mica crystal. An identical von Hámos spectrograph, equipped with Muscovite mica from the same cut, was also used to record the spontaneous emission emitted transversely with respect to the channel axis. Not shown is the location of a film pack used for measurement of the amplified x-ray beam composed of a 2 cm square 12.7 μm thick Ti foil backed by a matching piece of x-ray film. With removal of the axial von Hámos spectrograph, this detector was placed on the channel axis in a perpendicular orientation at a distance of 12.5 cm from the cluster target.

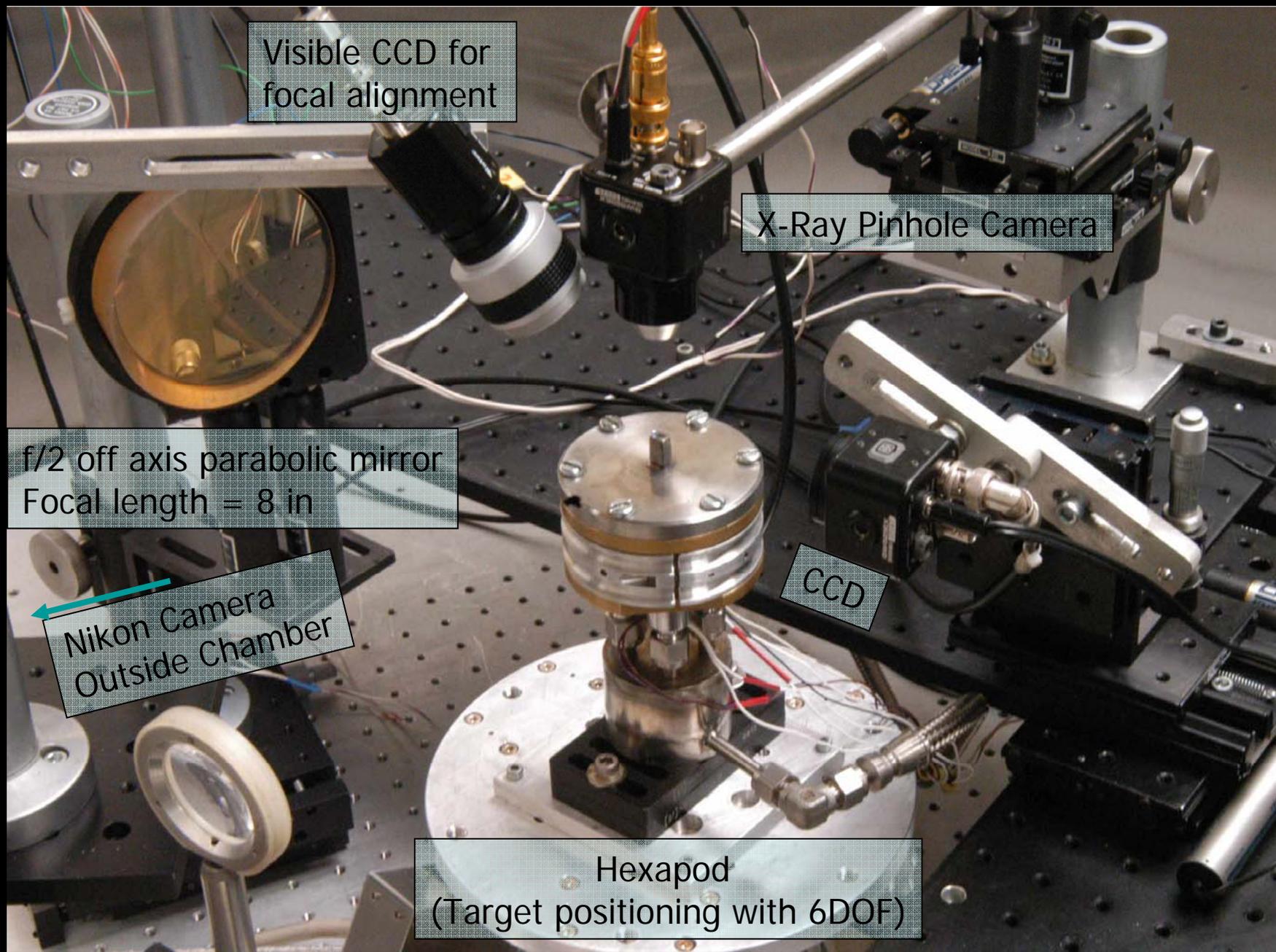
Input Variable

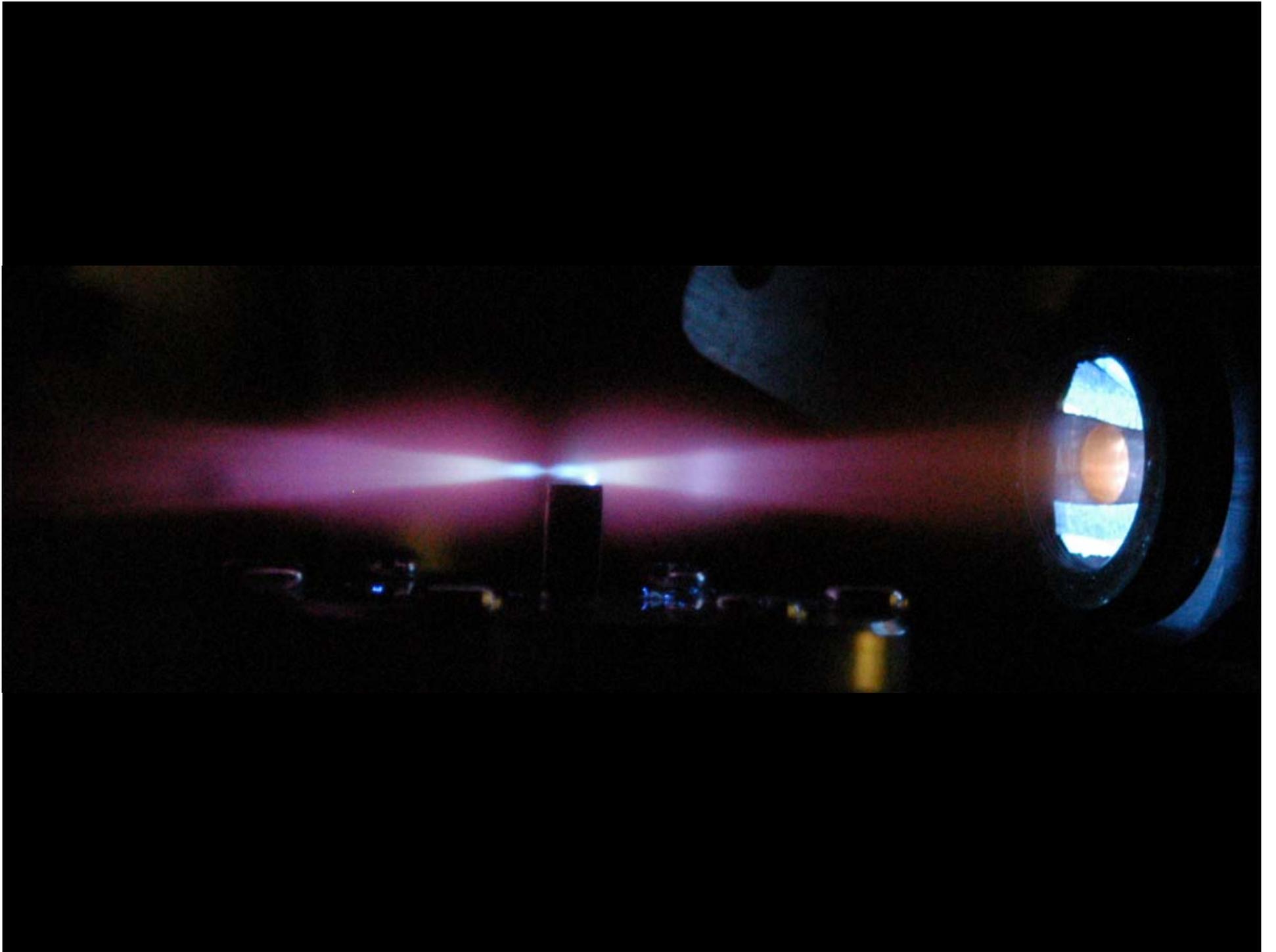
# Chamber Setup

Experimental Data Source

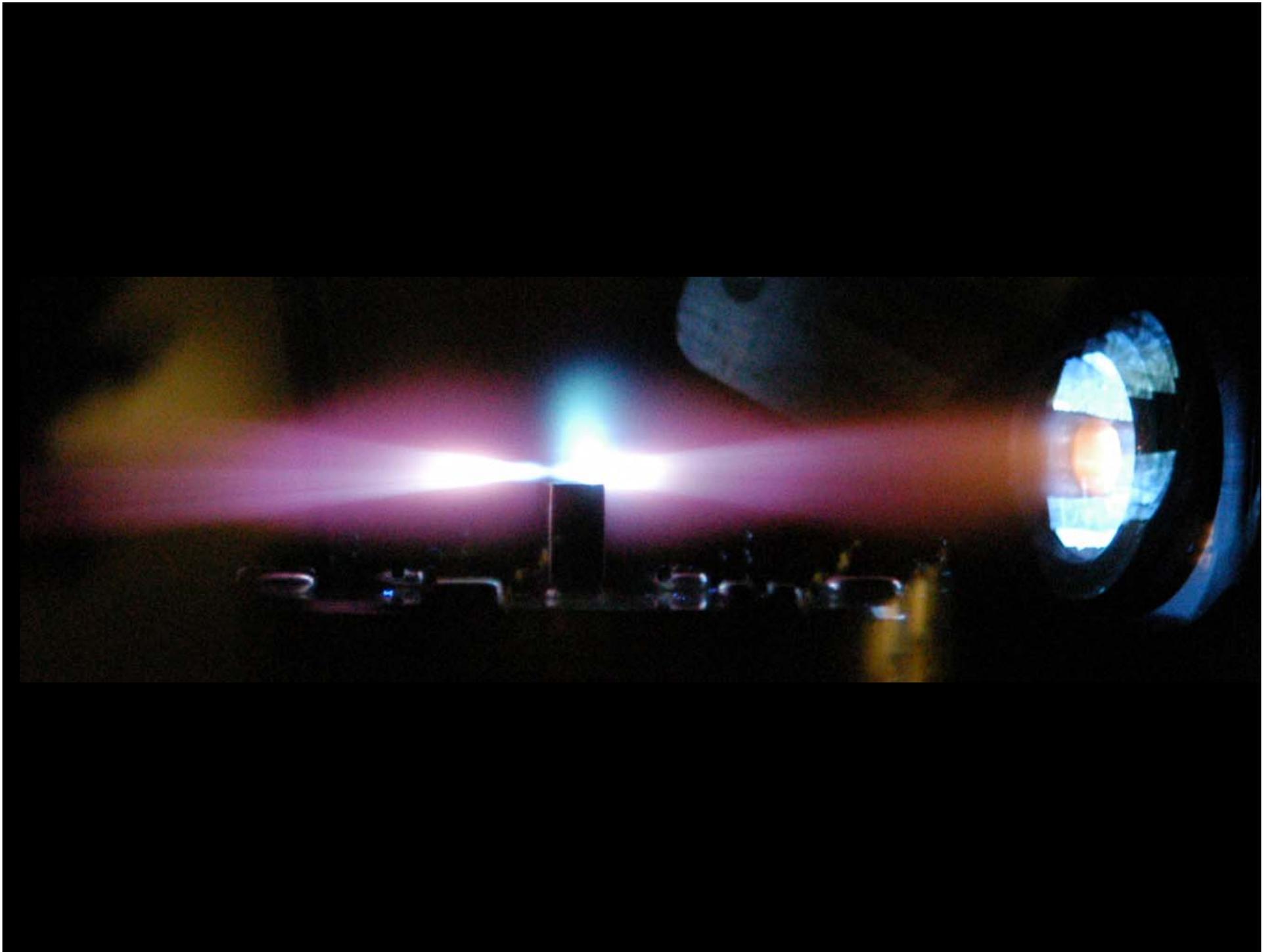


# Experimental Setup



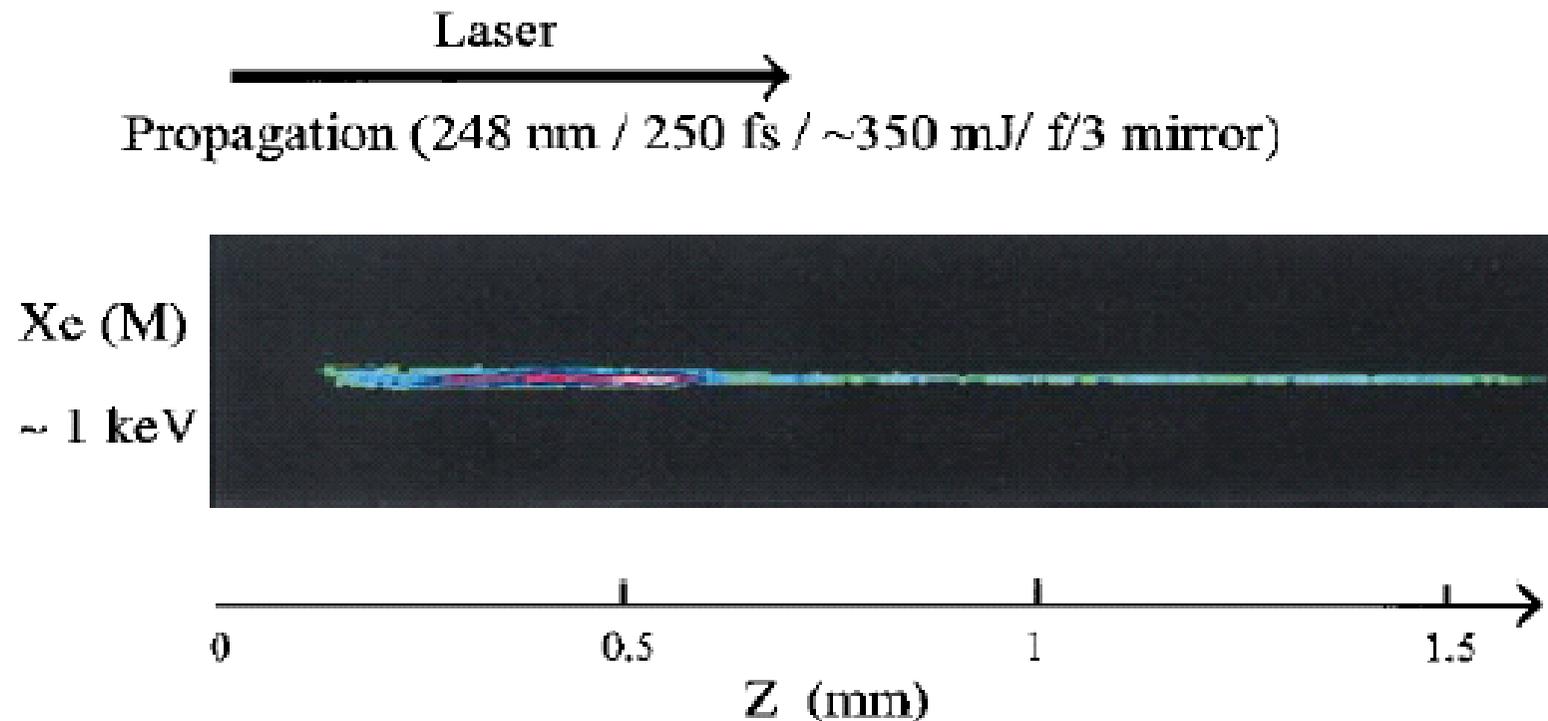






# CHANNEL MORPHOLOGY

## Xe(M) ~ 1 keV IMAGE



# Xe(L) SPONTANEOUS EMISSION SPECTRUM

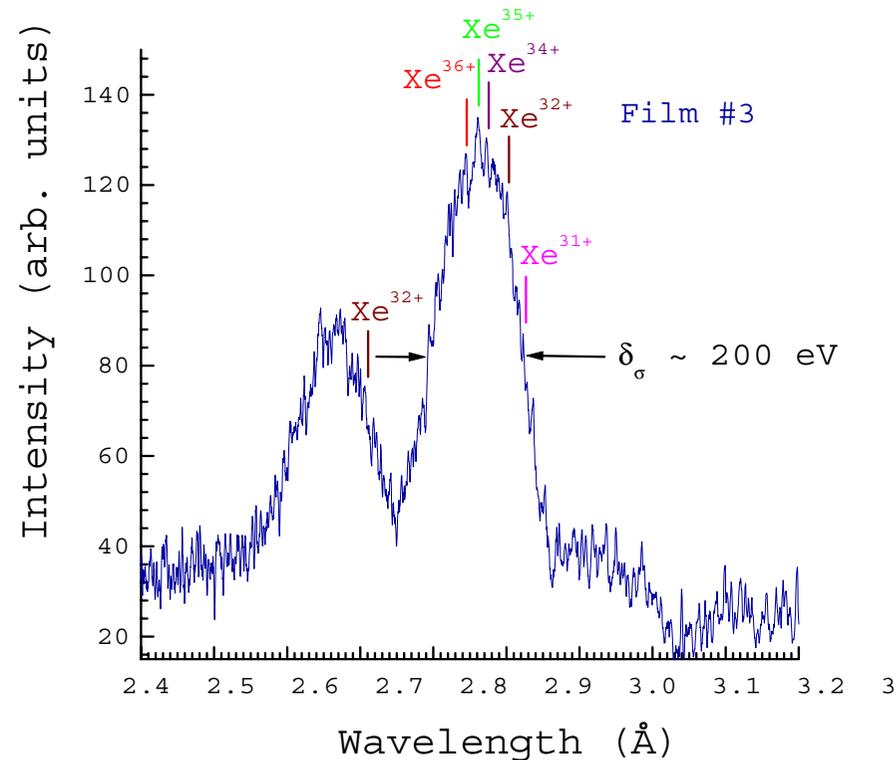
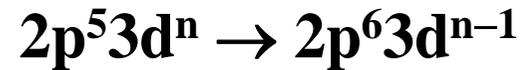


Fig. (1). Unamplified spontaneous emission profile of the Xe(L) 3d–2p hollow atom [7] spectrum (film #3) produced from Xe clusters with femtosecond 248 nm excitation without plasma channel formation. The splitting between the major and minor lobes arises from the spin-orbit interaction of the 2p vacancy. The full width of the main feature is  $\delta_\sigma \sim 200$  eV. The positions of selected charge state transition arrays ( $\text{Xe}^{31+}$ ,  $\text{Xe}^{32+}$ ,  $\text{Xe}^{34+}$ ,  $\text{Xe}^{35+}$ , and  $\text{Xe}^{36+}$ ) are indicated.

Xe(L)

# Xe(L) Amplification

Ruby

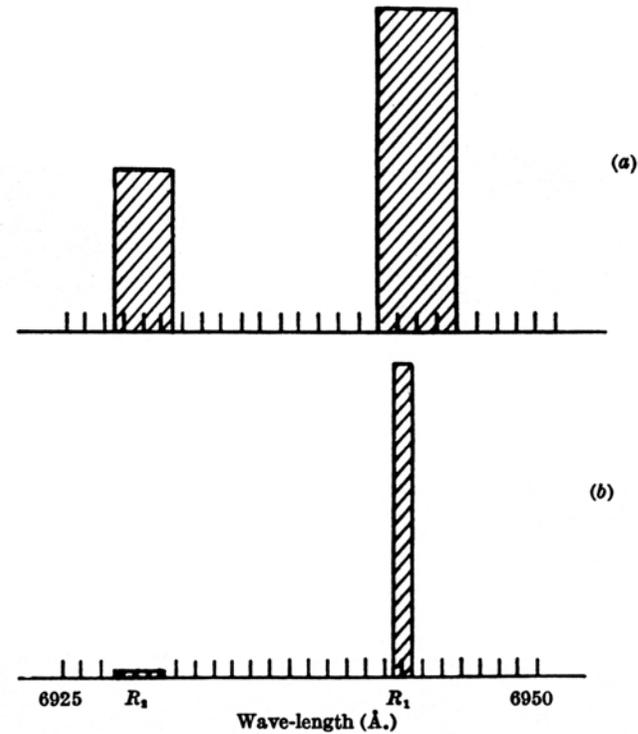
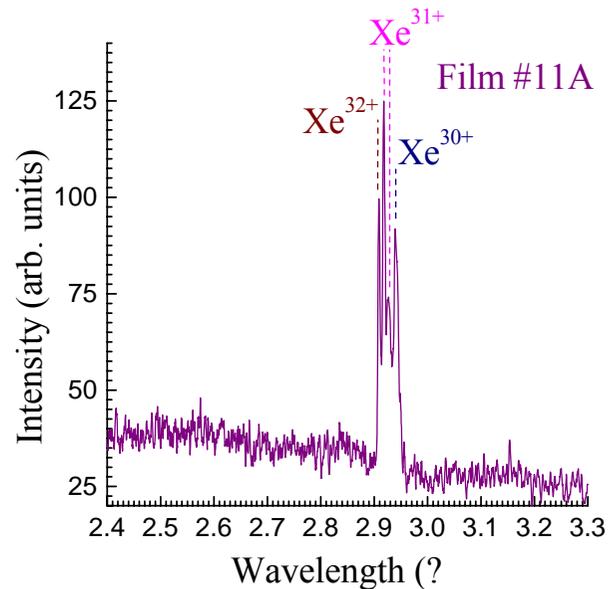
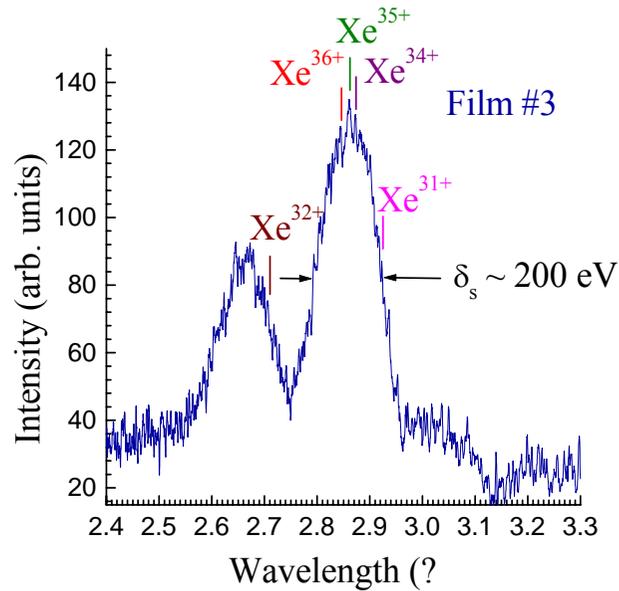


Fig. 2. Emission spectrum of ruby : *a*, low-power excitation ;  
*b*, high-power excitation

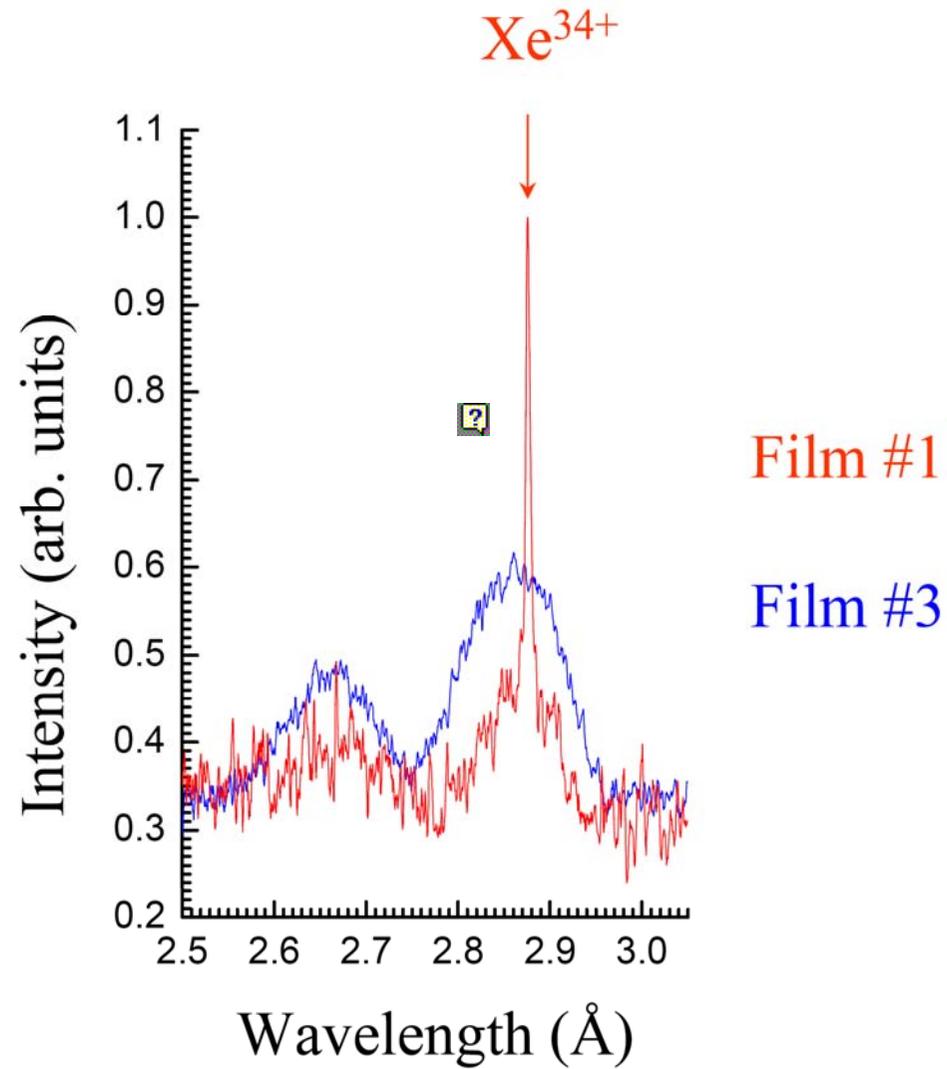
the emission spectrum obtained under these conditions is shown in Fig. 2*b*. These results can be explained on the basis that negative temperatures were produced and regenerative amplification ensued. I expect, in principle, a considerably greater ( $\sim 10^8$ ) reduction in line width when mode selection techniques are used<sup>1</sup>.

I gratefully acknowledge helpful discussions with G. Birnbaum, R. W. Hellwarth, L. C. Levitt, and R. A. Satten and am indebted to I. J. D'Haenens and C. K. Asawa for technical assistance in obtaining the measurements.

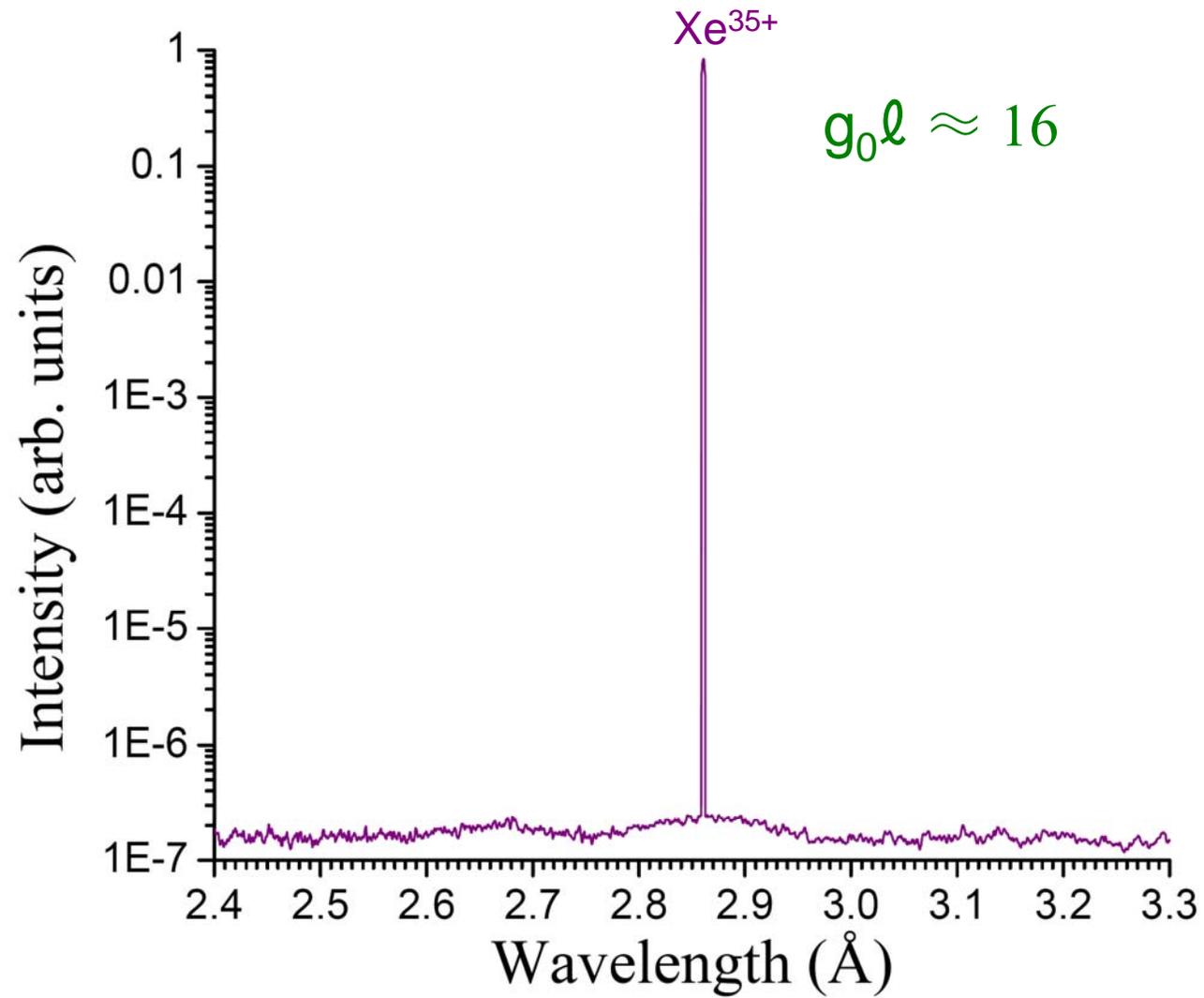
T. H. MAIMAN

Hughes Research Laboratories,  
A Division of Hughes Aircraft Co.,  
Malibu, California.

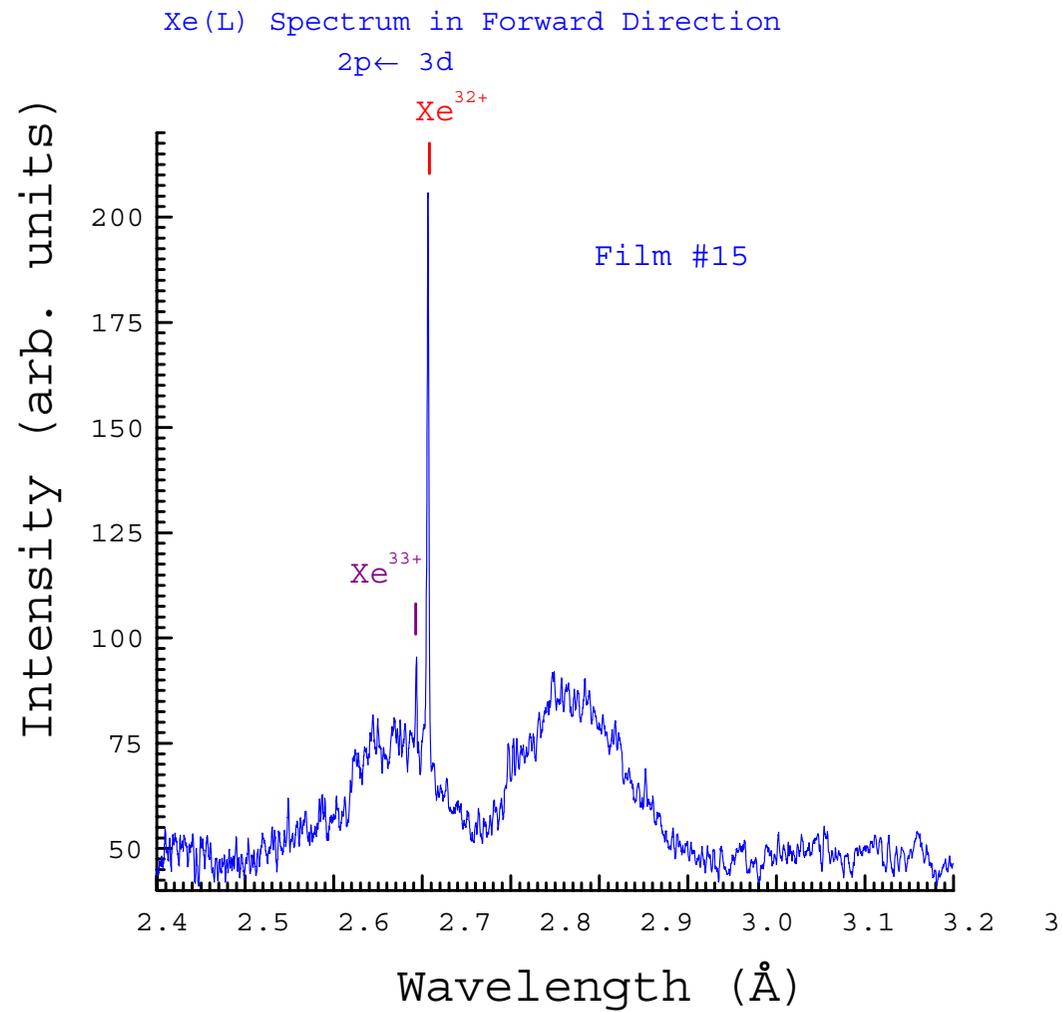
# Xe<sup>34+</sup> Line Amplification



# Amplified X-Ray Spectrum



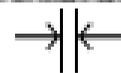
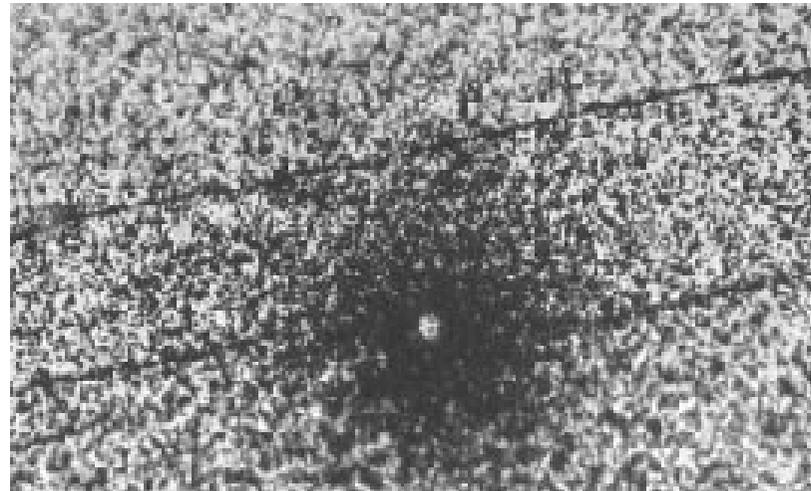
# Xe(L) AMPLIFICATION / Xe<sup>32+</sup> ~ 2.71 Å



# Xe<sup>32+</sup> (2.7 Å) Film Focus

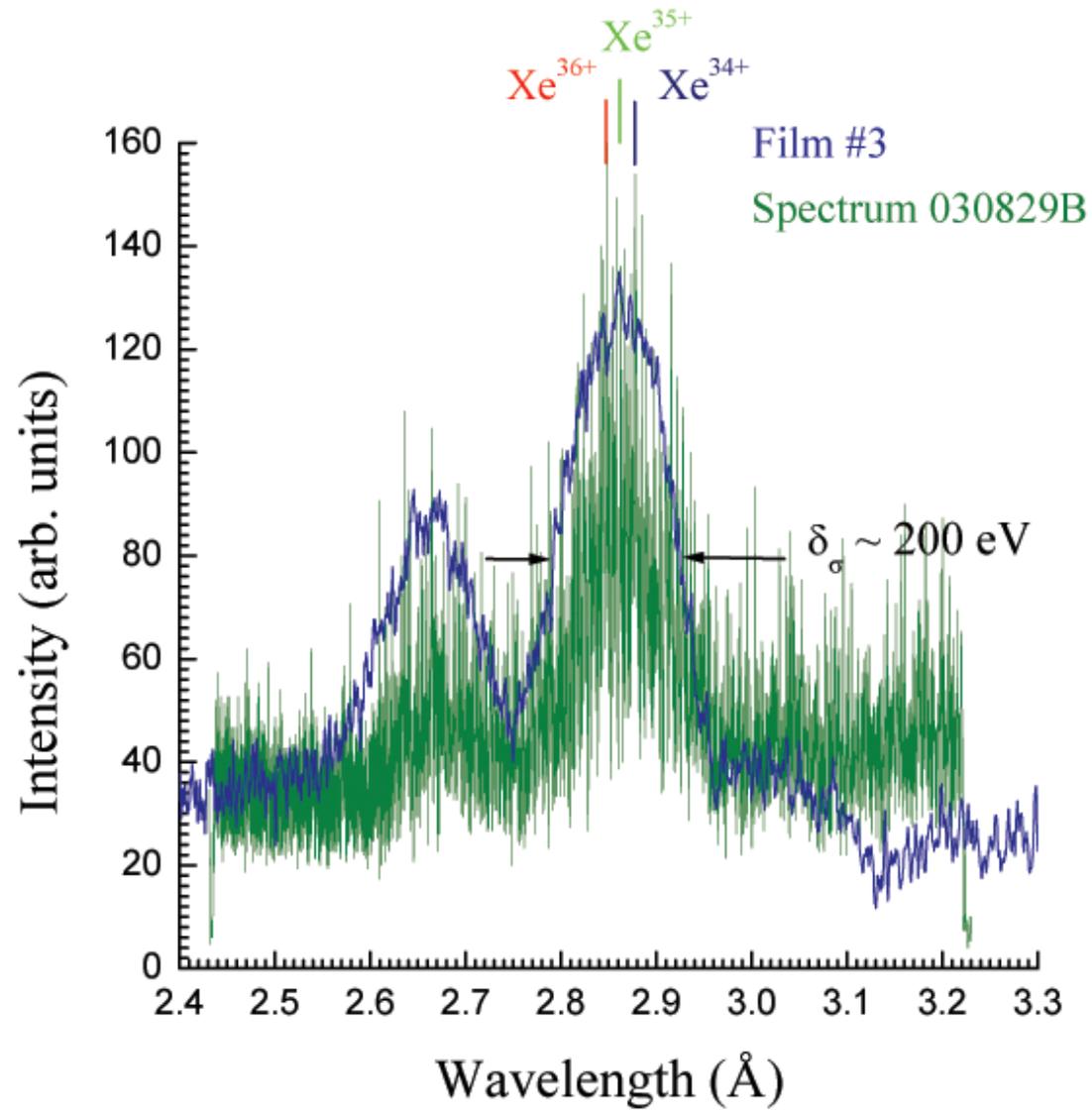
Film #15

$d_1 \cong 60 \mu\text{m}$



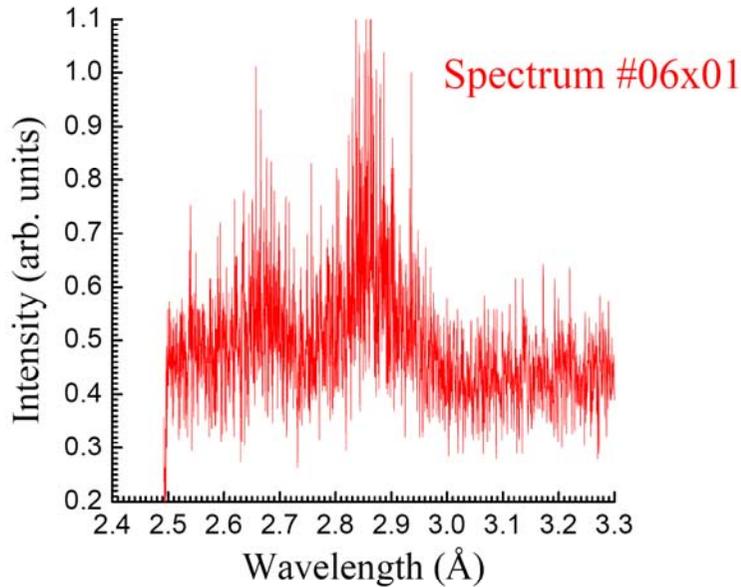
$d_2 \cong 5 \mu\text{m}$

# Single-Pulse Observation of Xe(L) Spectrum in Transverse Direction

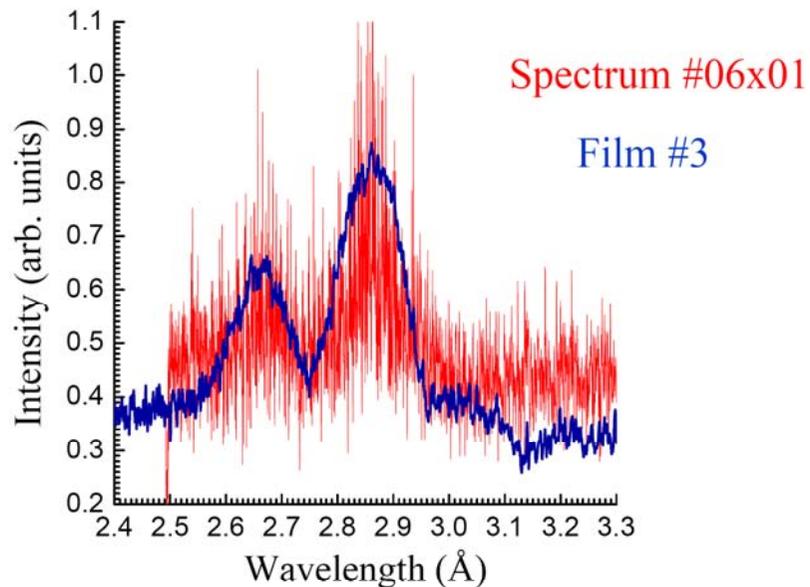


# Characteristic Unamplified **Single-Pulse** and **Time-Integrated** Spontaneous Emission Xe(L) Spectra

Characteristic Unamplified Axial **Single-Pulse** Xe(L) Spectrum



Characteristic Unamplified Axial **Single-Pulse** and Transverse **Time-Integrated** Xe(L) Spectra



Laser Propagation  
→

Xe (M)  
~ 1 keV



Propagation Image #06x01  
**No Channel Formation**

# SINGLE-PULSE (CCD) CONFIRMATION OF SPECTRAL HOLE-BURNING

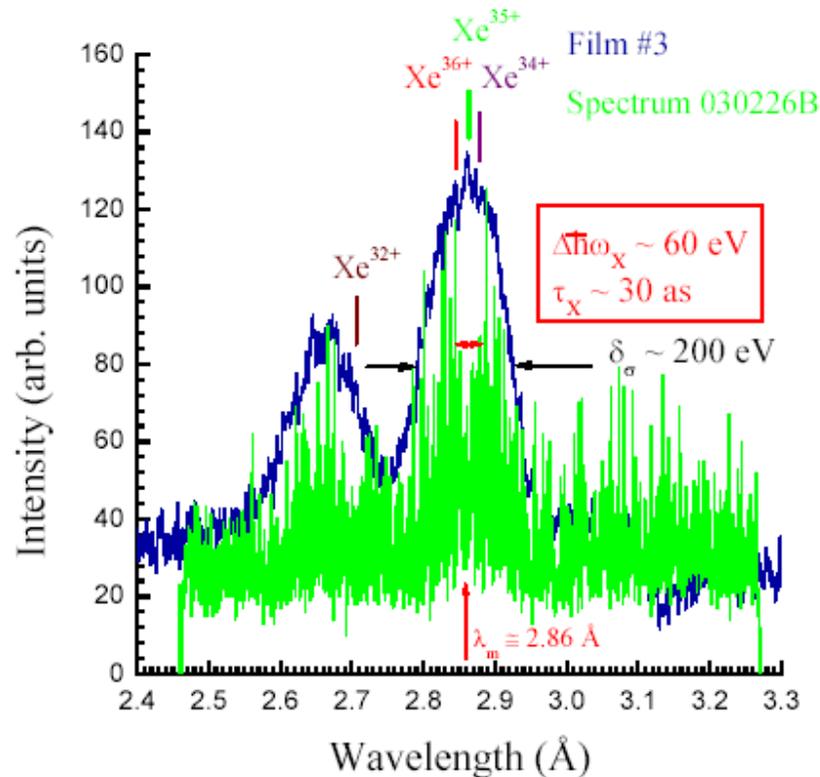
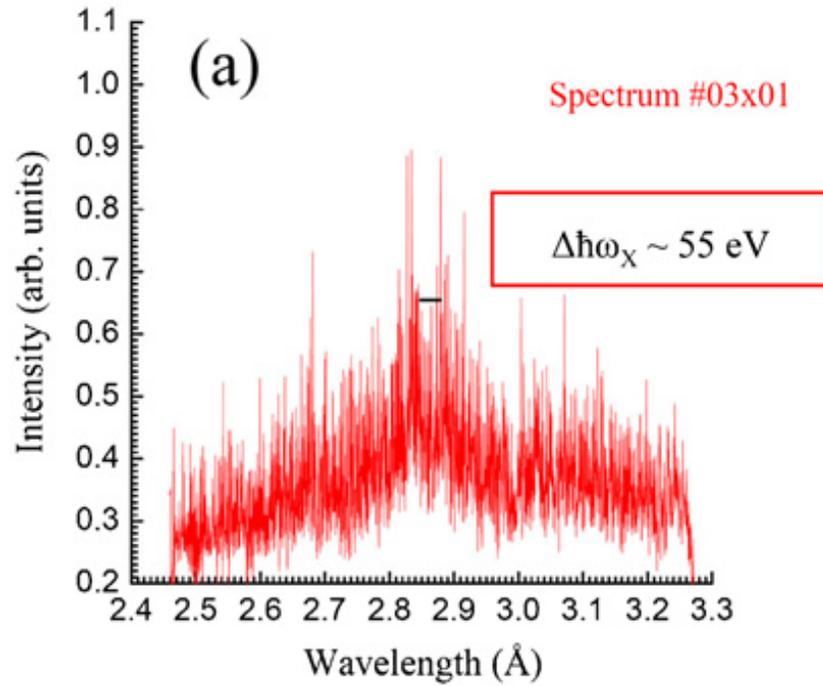


Fig. (3): Comparison of the characteristic spontaneous Xe(L) hollow atom emission spectrum (film #3) shown in Fig. (1) with a corresponding single-pulse spectrum (#030226B) recorded in a direction transverse to the channel. The principal feature of this comparison is the presence of a deep and broad ( $\Delta\hbar\omega_x \sim 60$  eV) spectrally hole-burned gap that matches well the location of the  $\text{Xe}^{34+}$  and  $\text{Xe}^{35+}$  arrays and is centered at  $\lambda_m \cong 2.86$  Å, the wavelength at which complete suppression of the transversely radiated emission is observed. The width  $\Delta\hbar\omega_x$  corresponds to a bandwidth sufficient for the amplification of multikilovolt x-ray pulses down to a limiting width  $\tau_x \sim 30$  as. An additional zone of spectral hole-burning is apparent near  $\lambda \cong 2.7$  Å. This region is associated with the  $\text{Xe}^{32+}$  array on the minor lobe that exhibited very strong amplification in the earlier study [8]. The spectral resolution of the CCD recorded data is estimated to be  $\sim 4$  eV.

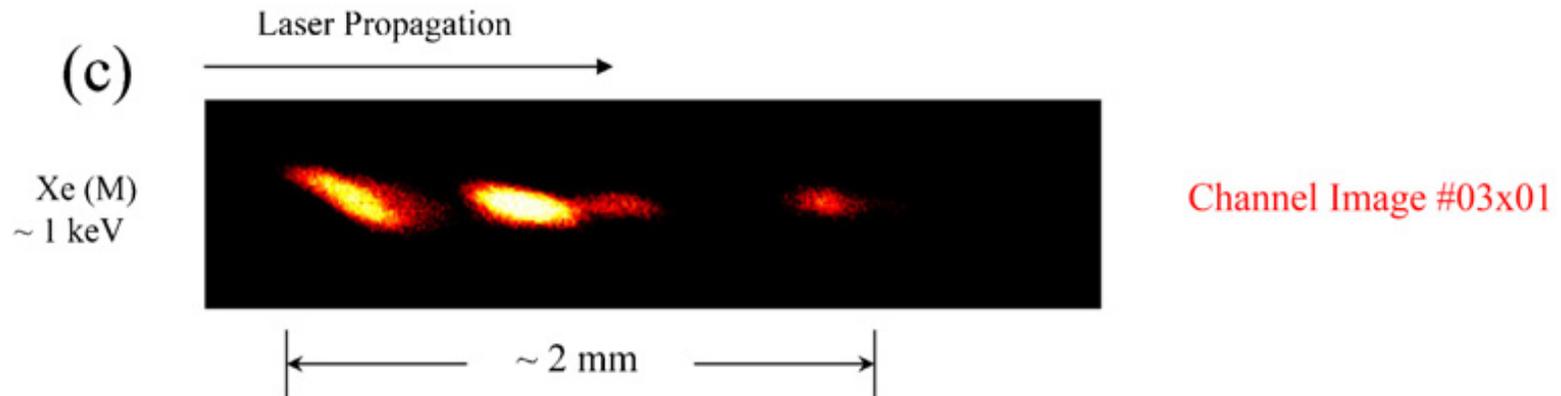
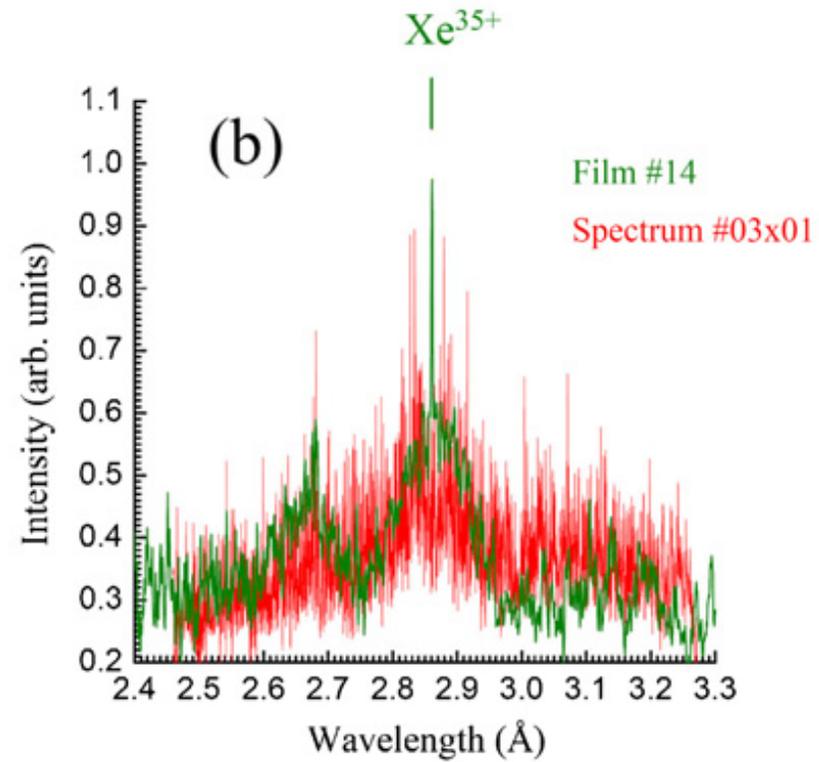
- Full Signal Extinction at  $\lambda_m \sim 2.86$  Å ( $\text{Xe}^{35+}$ )
- Spectral Width  $\Delta\hbar\omega_x \sim 60$  eV  $\Rightarrow \sim 30$  as  
(Atomic Time  $\hbar/m_e c^2 \sim 24$  as)

# Xe(L) Saturated Amplification Spectral Width Measurement

Transverse Single-Pulse  
Xe(L) Spectrum

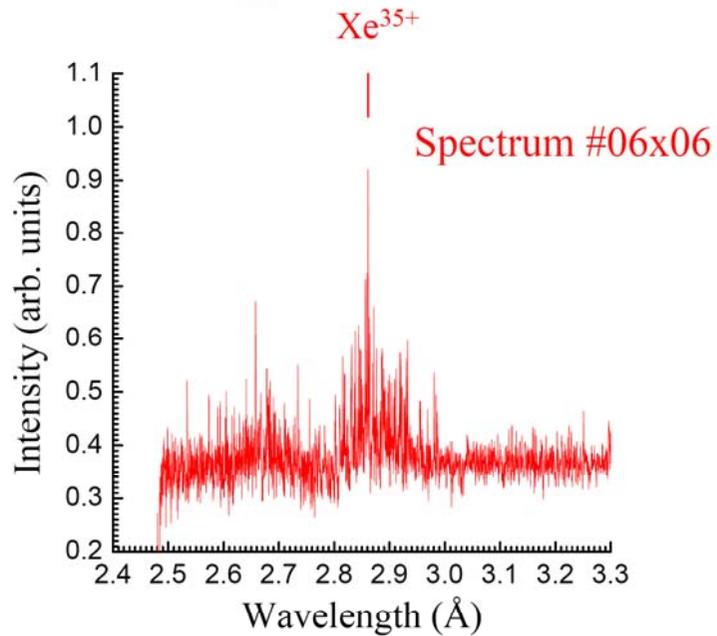


Xe(L) Spectral Hole Burning

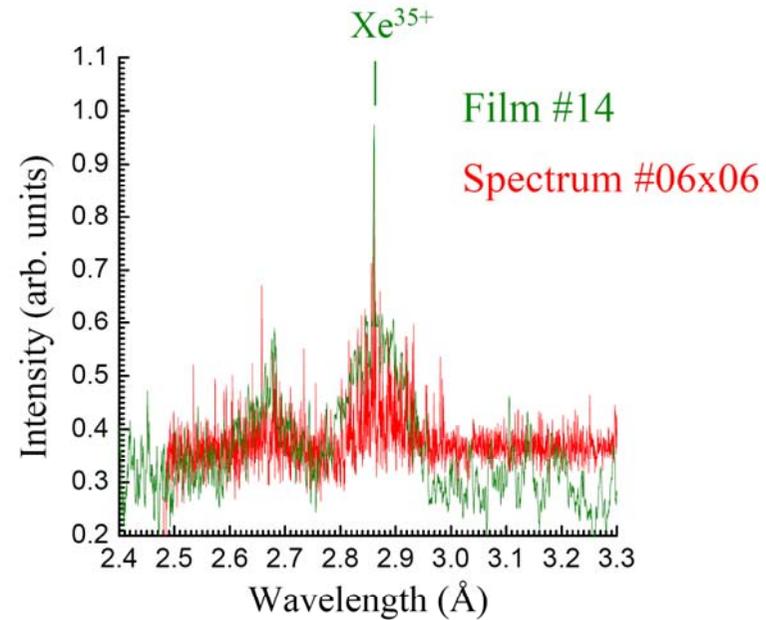


# Axial Amplified Single-Pulse and Time-Integrated Xe(L) Spectra

Axial **Single-Pulse** Xe(L) Spectrum

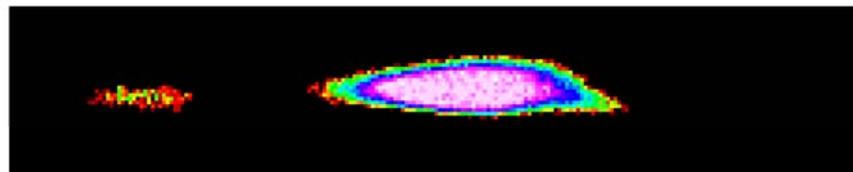


Axial **Single-Pulse** and **Time-Integrated** Spectra



Laser Propagation  
→

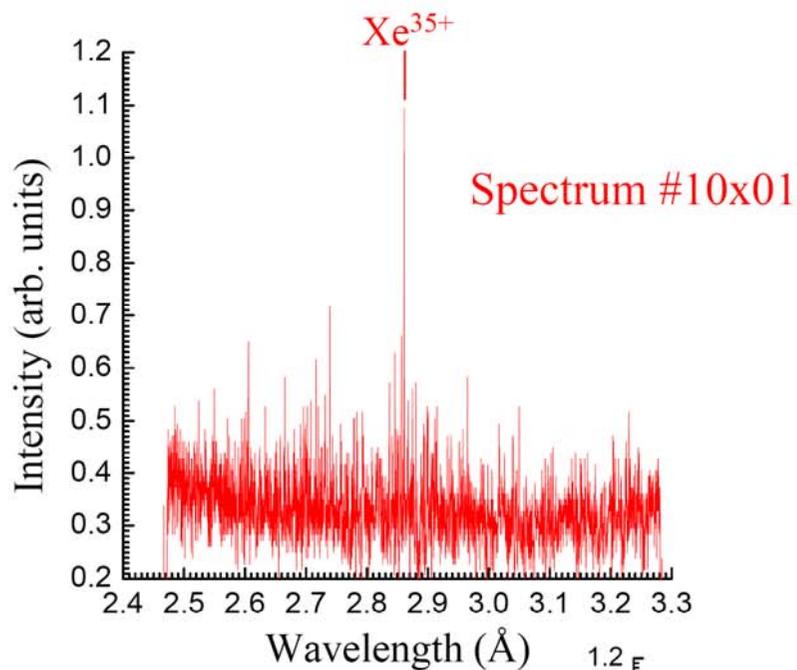
Xe (M)  
~ 1 keV



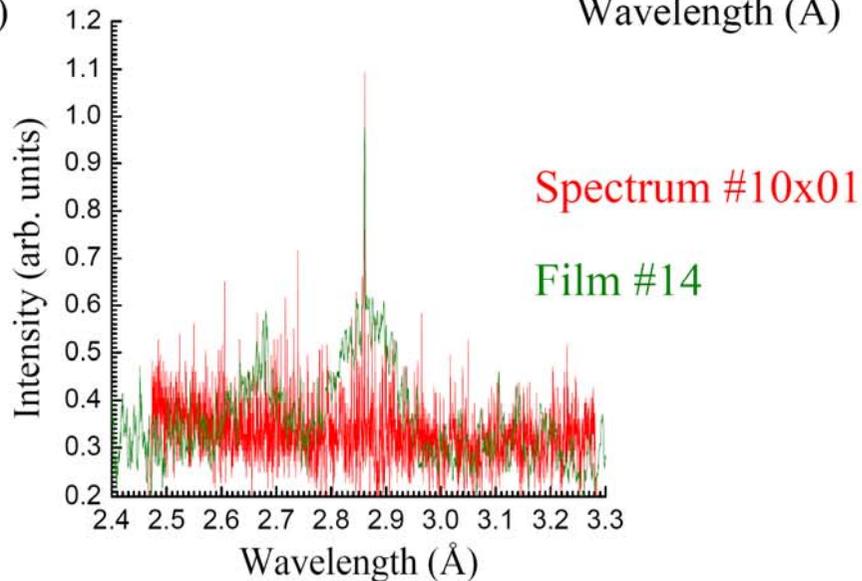
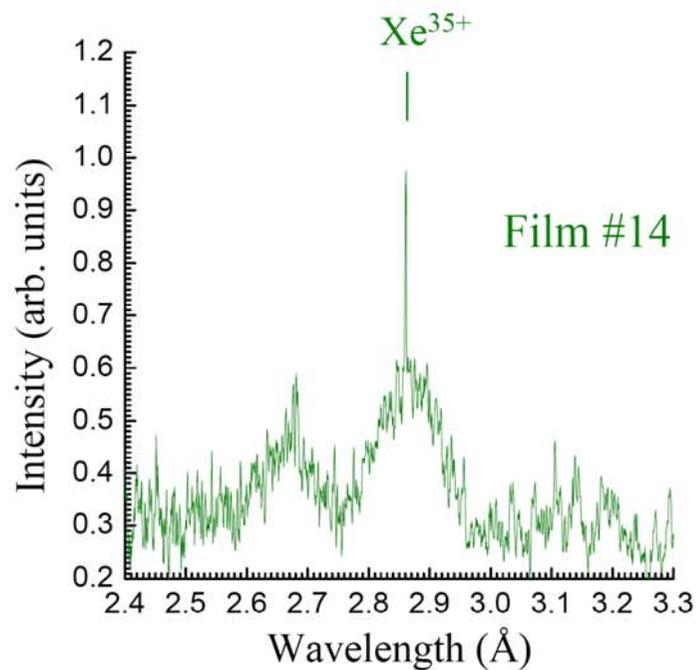
~ 1.8 mm

# Axial Amplified Single-Pulse and Time-Integrated Xe(L) Spectra

Axial **Single-Pulse** Xe(L) Spectrum

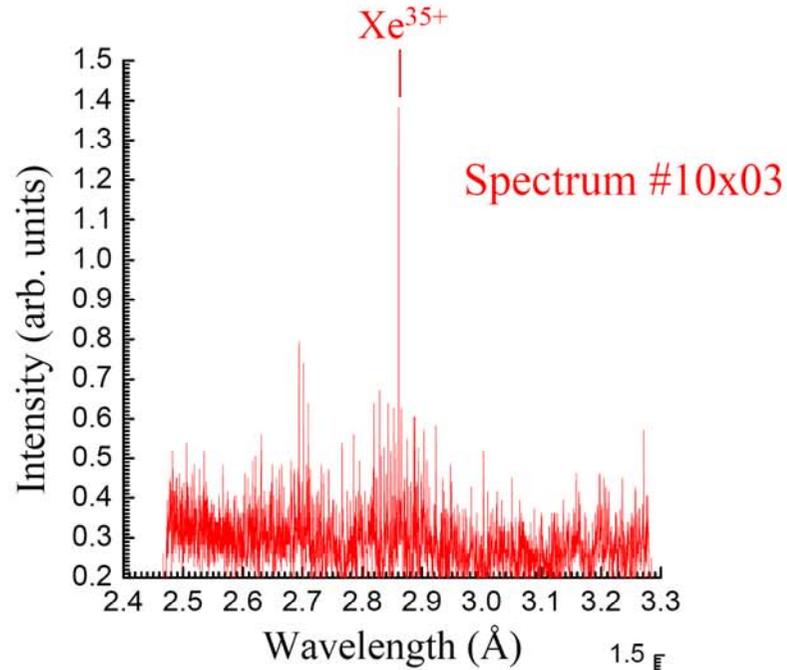


Axial **Time-Integrated** Spectrum

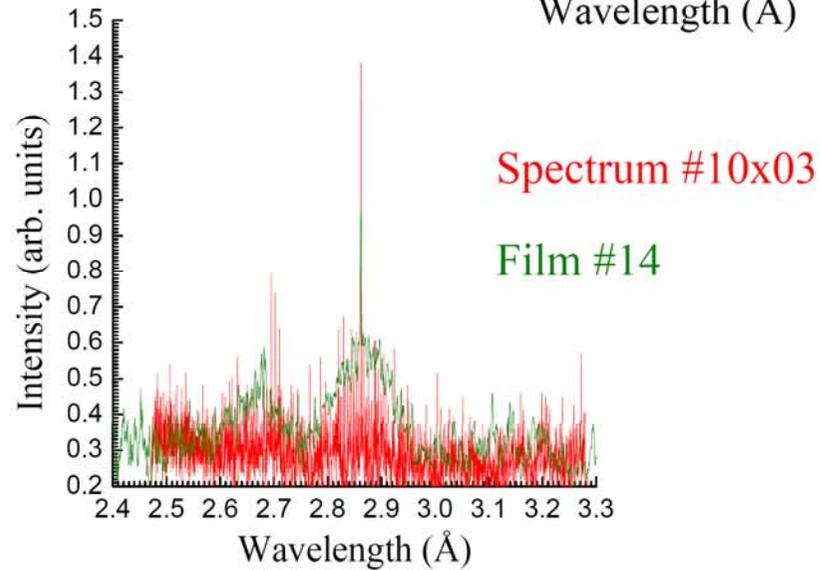
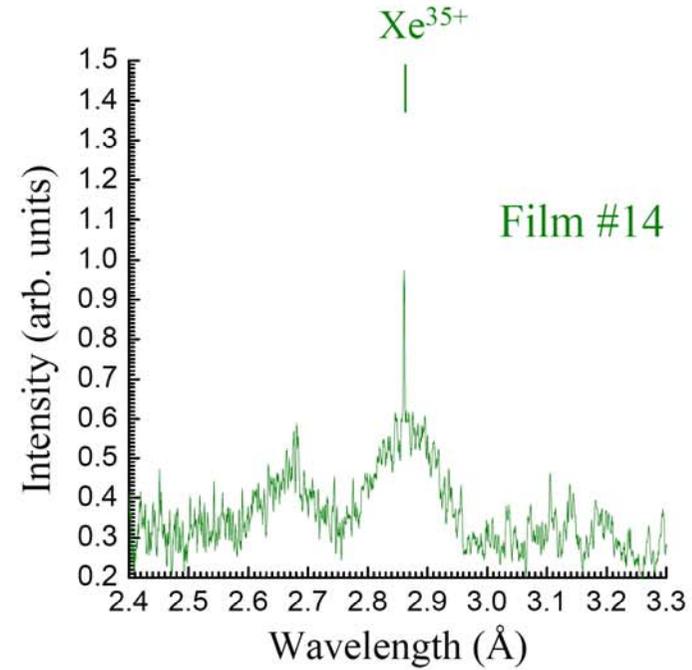


# Axial Amplified Single-Pulse and Time-Integrated Xe(L) Spectra

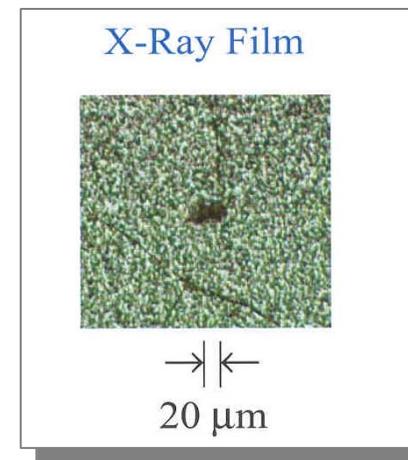
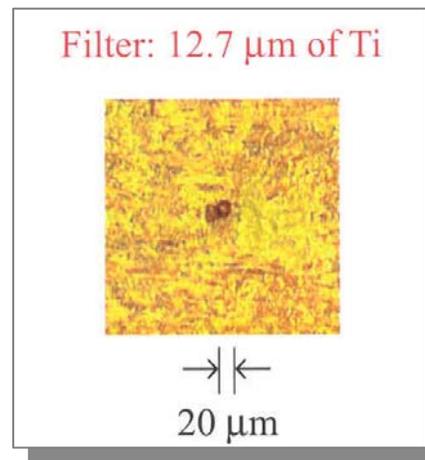
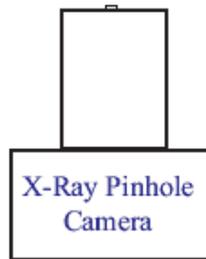
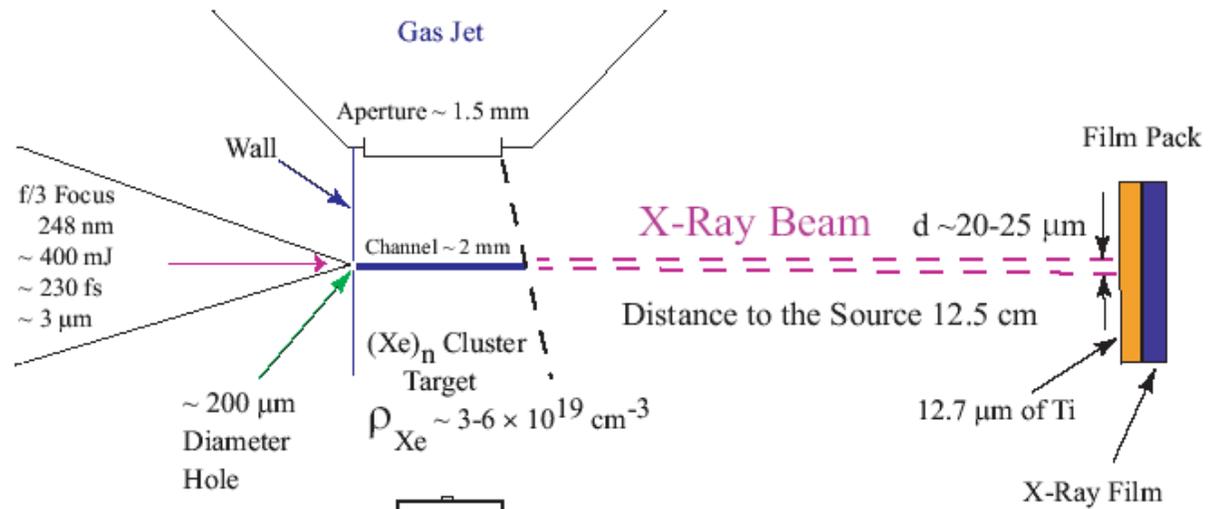
Axial **Single-Pulse** Xe(L) Spectrum



Axial **Time-Integrated** Spectrum

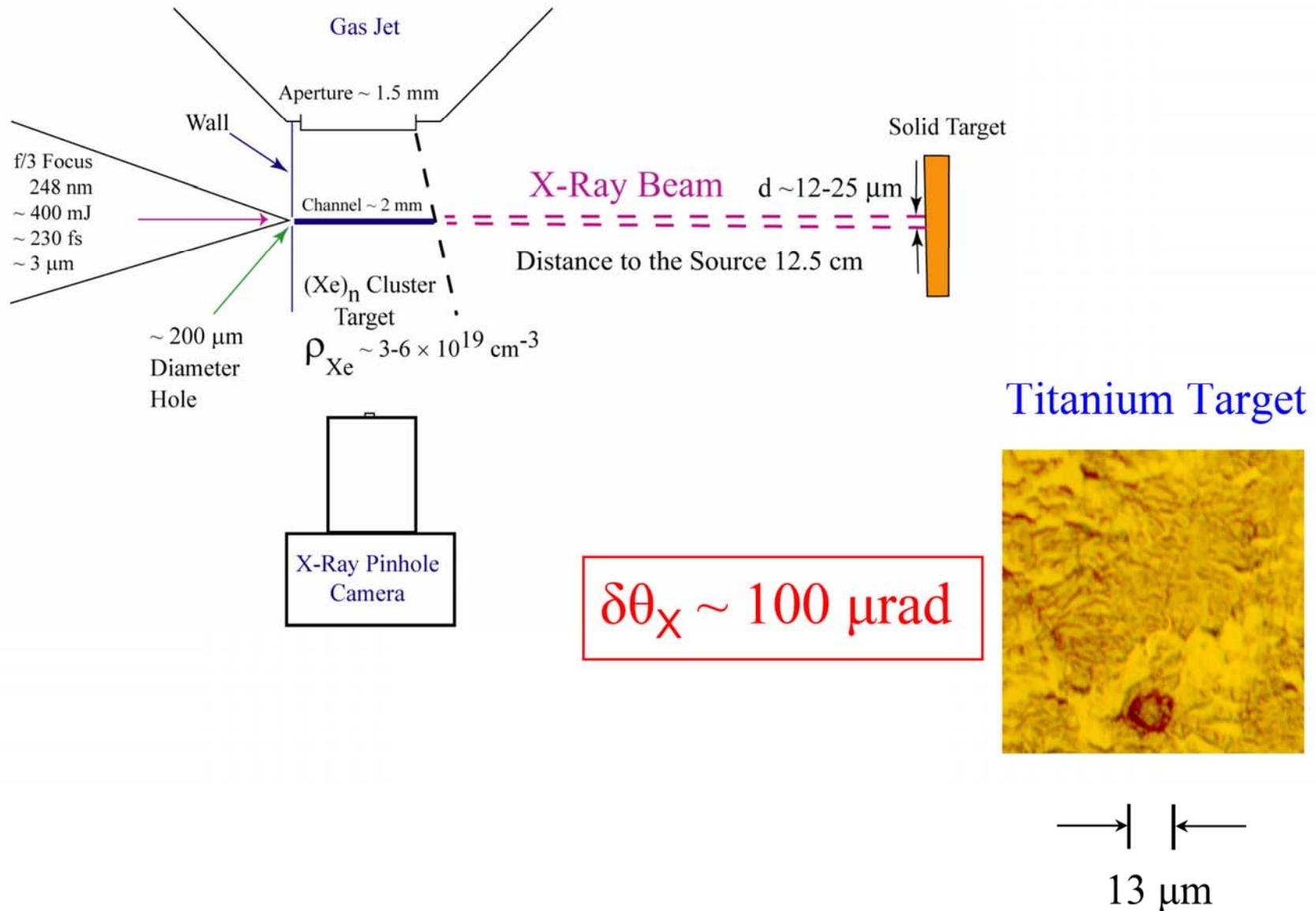


# X-RAY BEAM DOUBLE VISION

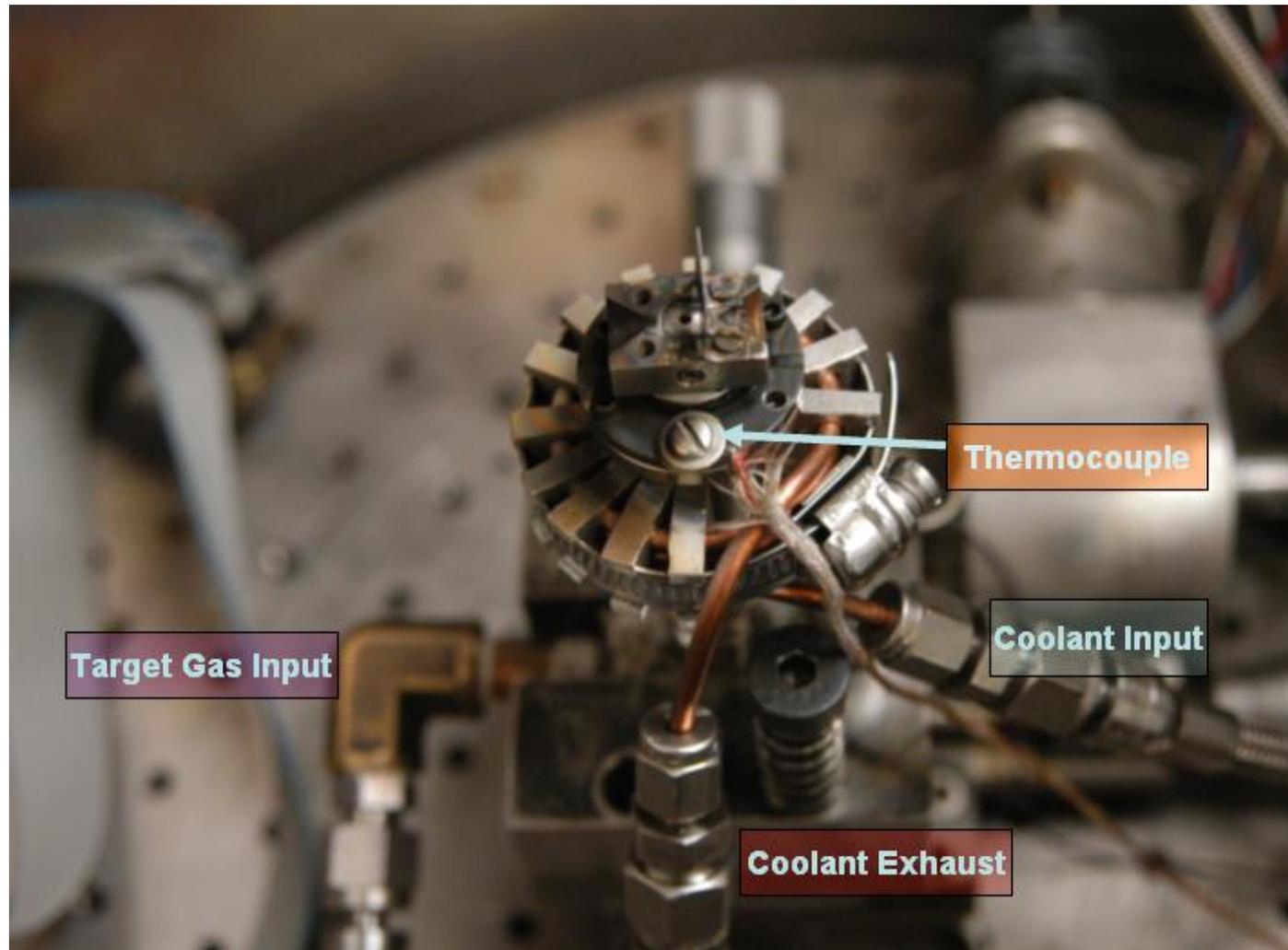


$$\delta\theta_x \sim 200 \mu\text{r}$$

# Experimental Data Xe(L) on Solid Targets



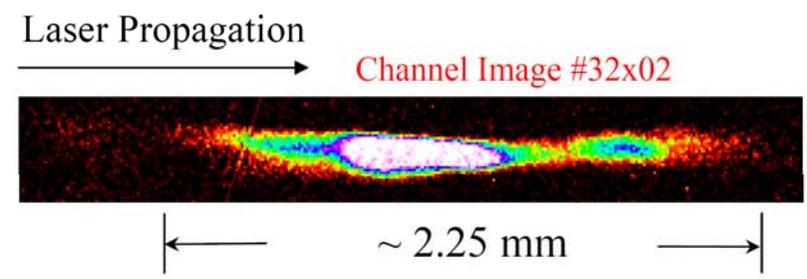
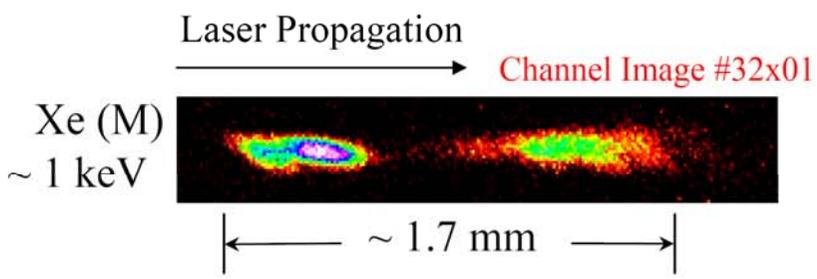
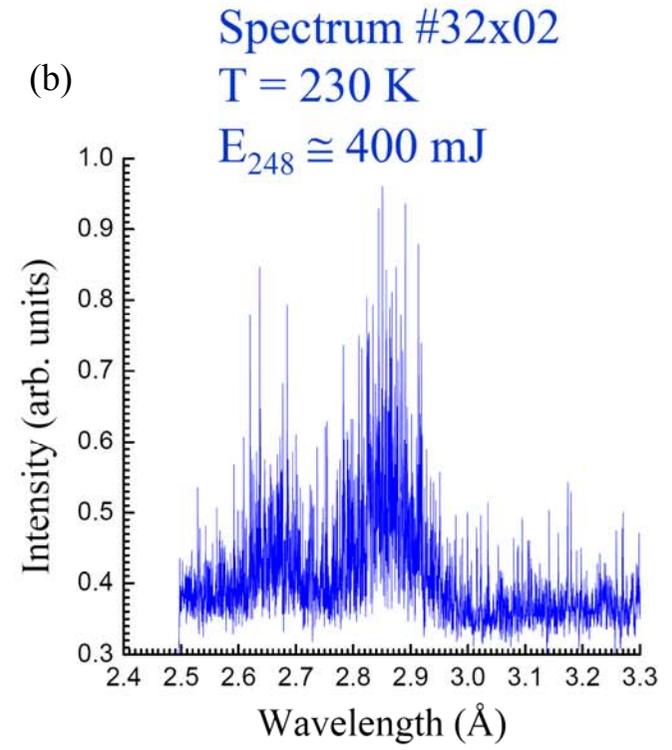
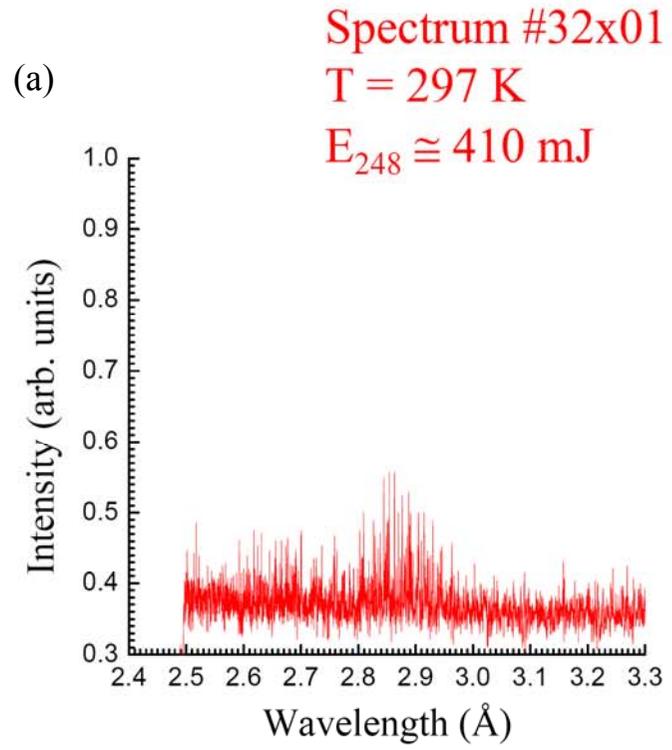
## COOLED CLUSTER FORMATION



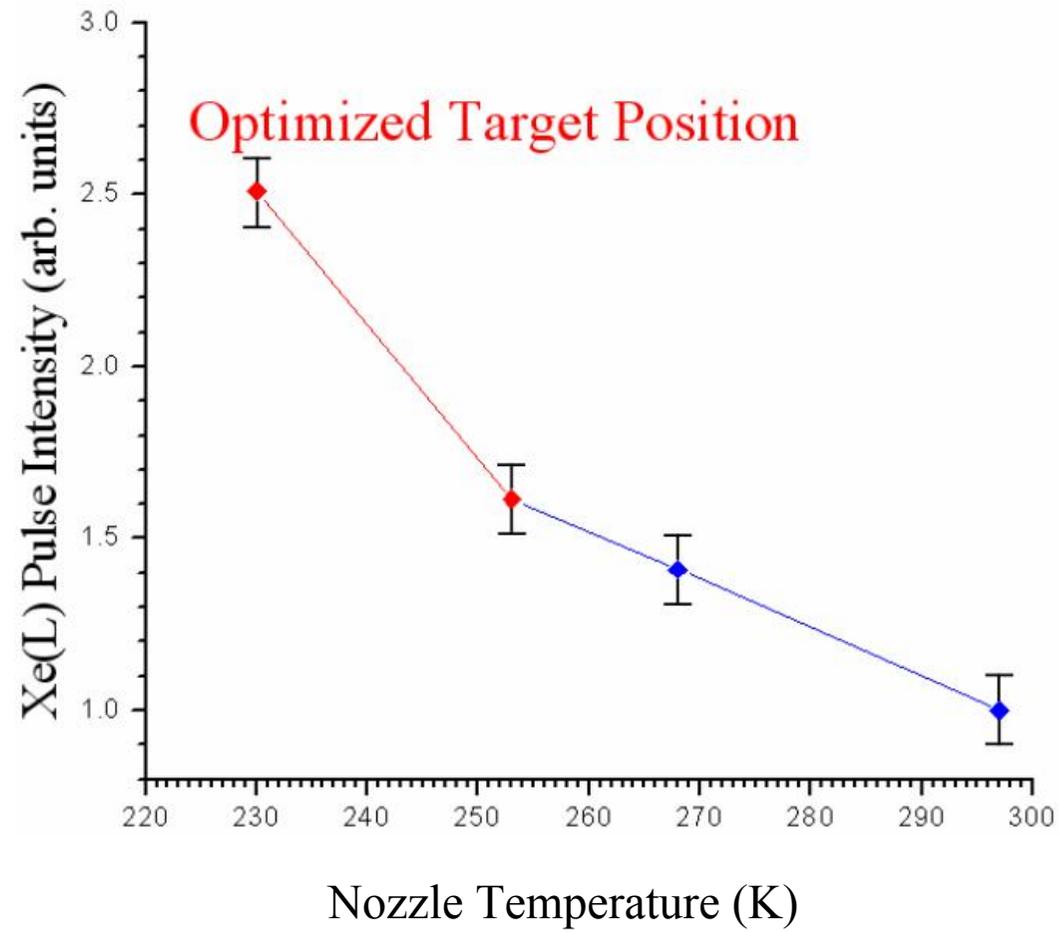
Picture of cooled Xe nozzle showing the coolant coil and the Xe target gas input. The temperature is measured with a thermocouple located on the body of the nozzle. The nozzle emits the xenon gas plume upward in this bird's-eye view.

# Temperature Dependence of Xe(L) X-Ray Pulse Intensity

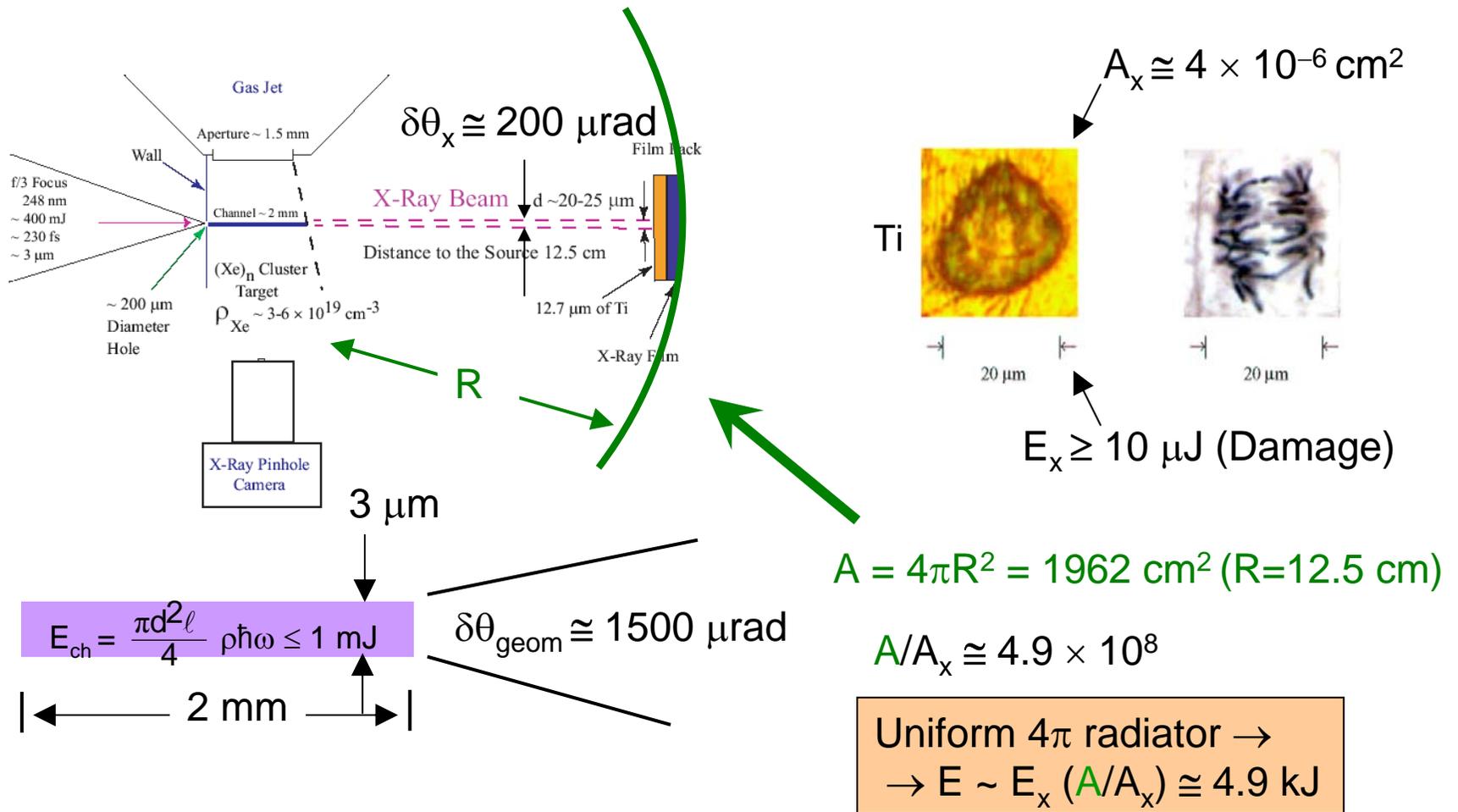
## Single-Pulse Xe(L) Spectra in Forward Direction



## Temperature Dependence of Xe(L) X-Ray Emission Intensity



# AMPLIFIER GAIN

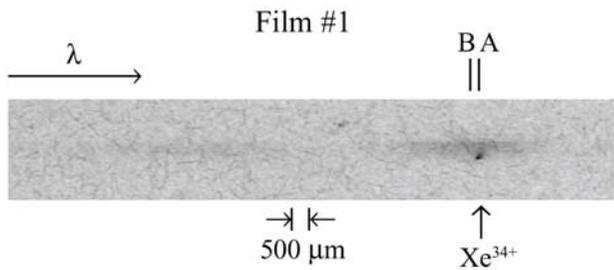
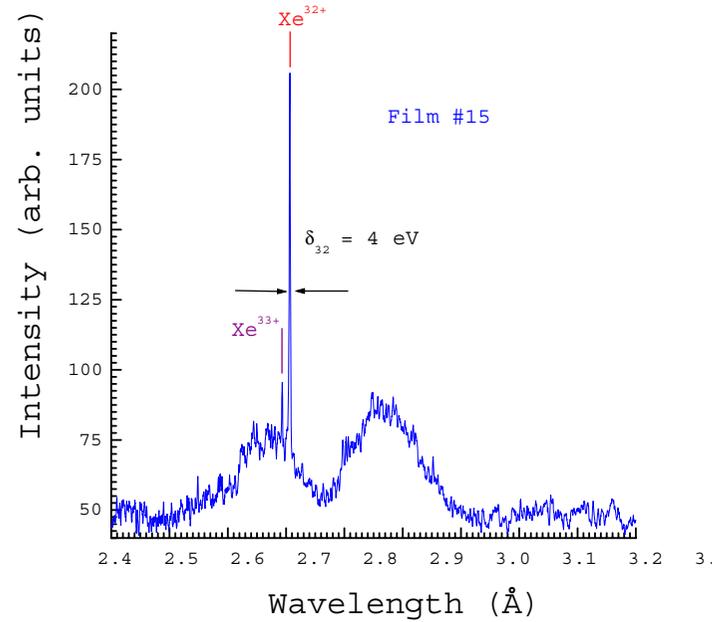
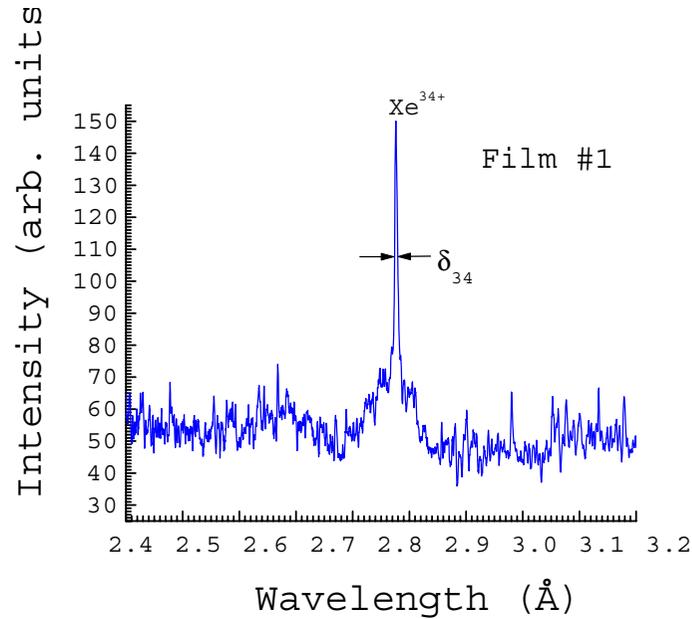


$$\delta\theta_{geom} \cong \frac{3 \times 10^{-4}}{2 \times 10^{-1}} = 1.5 \times 10^{-3} = 1500 \mu\text{rad} \gg \delta\theta_x = 200 \mu\text{rad}$$

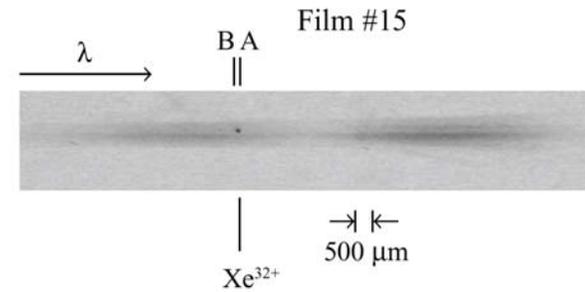
← measured

$$\text{Gain} = G \sim E/E_{ch} \cong \frac{4.9 \text{ kJ}}{1 \text{ mJ}} \cong 5 \times 10^6 \sim 1-5 \times 10^6$$

# GAIN COMPARISONS



$$G \sim 10^6$$



$$G \sim 6 \times 10^6$$

*J. Phys. B* **36**, 3433 (2003)

From Damage (Ti foil)

$$G \sim 1-5 \times 10^6$$

# CHANNEL STABILITY, EIGENMODE, AND CRITICAL POWER

Physics: Borisov *et al.*

Proc. Natl. Acad. Sci. USA 95 (1998) 7855

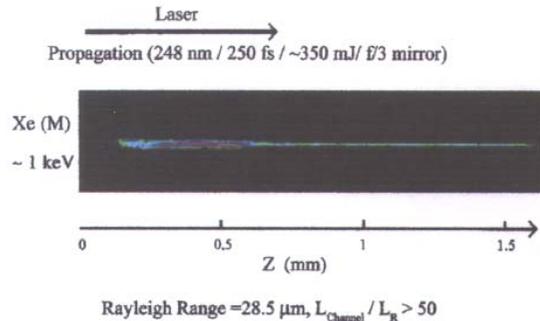


FIG. 1. Single-exposure x-ray image of a stable slender channel emitting Xe(M) radiation ( $\sim 1$  keV) produced in a gaseous target containing  $(\text{Xe})_0$  clusters. The incident 248-nm pulse had a duration of  $\sim 250$  fs and an energy of  $\sim 350$  mJ, and was focused with  $f/3$  off-axis parabolic mirror. The image was recorded with an x-ray pinhole camera having an aperture with a diameter of  $25 \mu\text{m}$  and a spatial resolution of  $\sim 30 \mu\text{m}$ . Because the Rayleigh length of the focusing system is  $28.5 \mu\text{m}$ , the observed length of the channel exceeds 50 Rayleigh ranges. Additional experimental details are reported in ref. 4. The color scale (in arbitrary units) of the measured x-ray intensity is defined by black, zero; red through violet, ascending intensity; and white, maximum.

limiting the stability of the confined propagation illustrated in Fig. 1, and the evaluation of the corresponding upper bound on the power density.

Following conventional notation (1), we introduce the definitions of the coordinates of the  $\eta$ - $\rho_0$  plane given by

$$\eta = P_0/P_{cr}, \quad \rho_0 = r_0\omega_p/c, \quad [1]$$

with  $P_0$  denoting the incident peak power and with the critical power ( $P_{cr}$ ) given by (1, 12)

$$P_{cr} = (m_{e,0}^2 c^5 / e^2) \int_0^\infty g_0^2(\rho) \rho d\rho (\omega/\omega_{p,0})^2 = 1.6198 \times 10^{10} (\omega/\omega_{p,0})^2 W, \quad [2]$$

in which  $m_{e,0}$ ,  $c$ , and  $e$  have their customary identifications  $g_0(\rho)$  is the Townes mode (21), and  $\omega$ ,  $\omega_{p,0}$ , and  $r_0$ , are the angular frequency corresponding to the propagating radiation the angular frequency of the unperturbed plasma, and the radius of the incident intensity profile, respectively. In addition, lowest eigenmodes (1, 10, 12) exist with the dimensionless radius

$$\rho_{e,0} = \left[ 2 \int_0^\infty U_{s,0}^2(\rho) \rho d\rho / U_{s,0}^2(0) \right]^{1/2}, \quad [3]$$

in which  $U_{s,0}(\rho)$  represents the eigenmode (1, 10, 12) with index  $s$ . The present analysis was confined to electron densities  $N_e$  less than one-quarter of the critical electron density ( $N_{cr}$ ) to eliminate resonant plasma wave production, and forward Raman scattering (22, 23) was not included, because it is known experimentally that it can be suppressed (24).

Fig. 2 illustrates the geography in the  $\eta$ - $\rho_0$  plane of the stable and unstable regions characteristic of channel formation in initially homogeneous plasmas (1). The essential features are the locus of the eigenmode curve  $\rho_{e,0}(\eta)$  defined by Eq. 3, the existence of a region of stable propagation that includes the

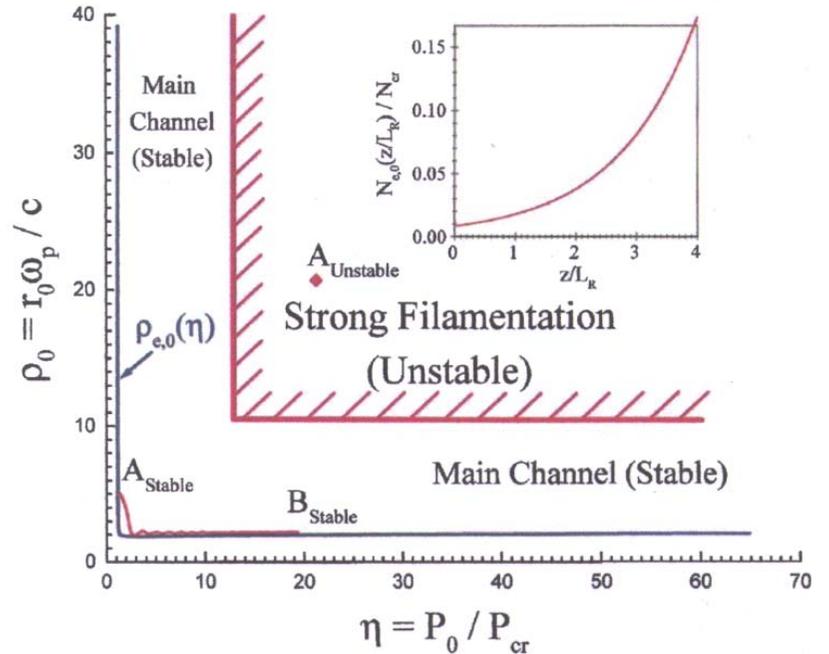
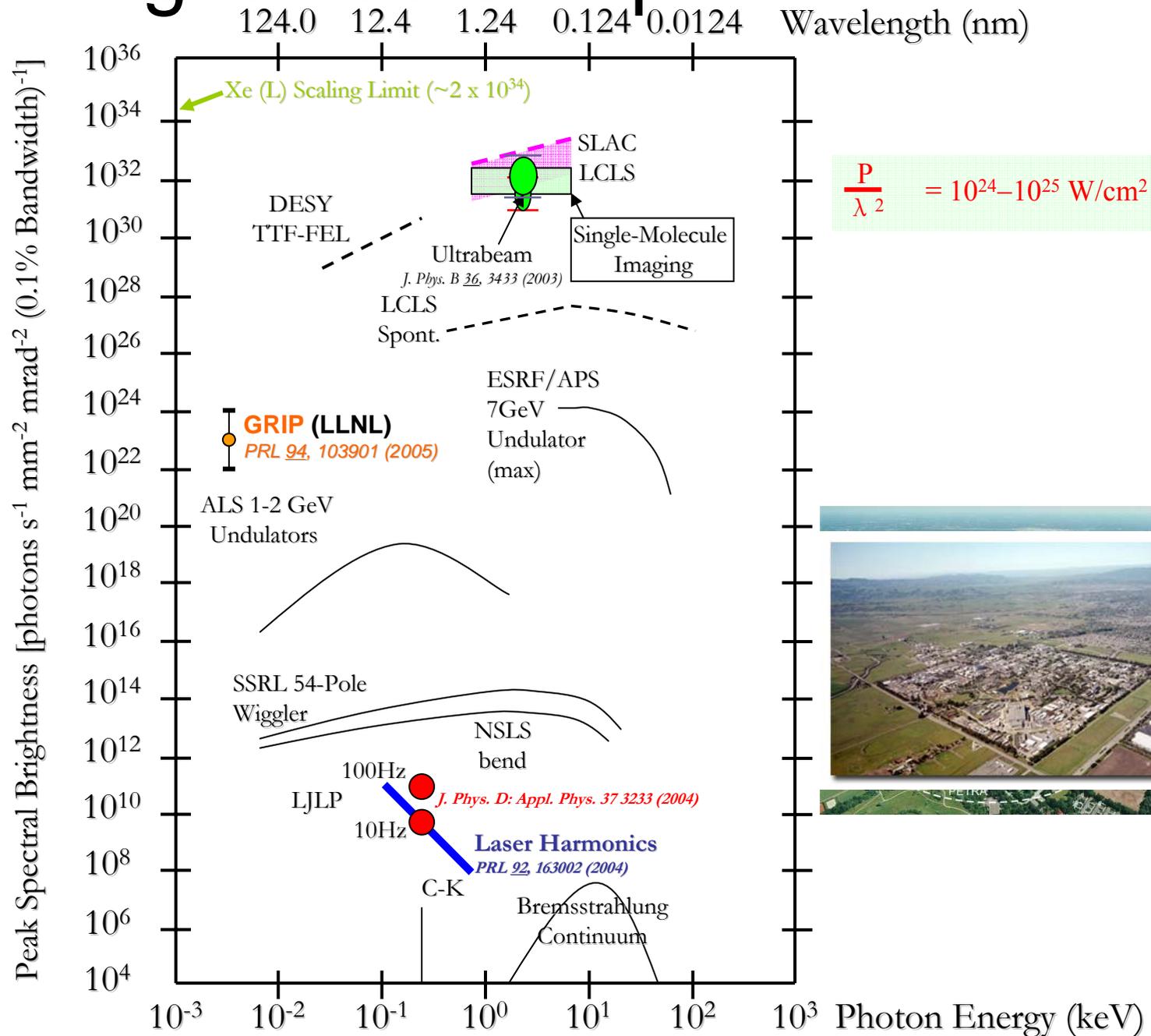
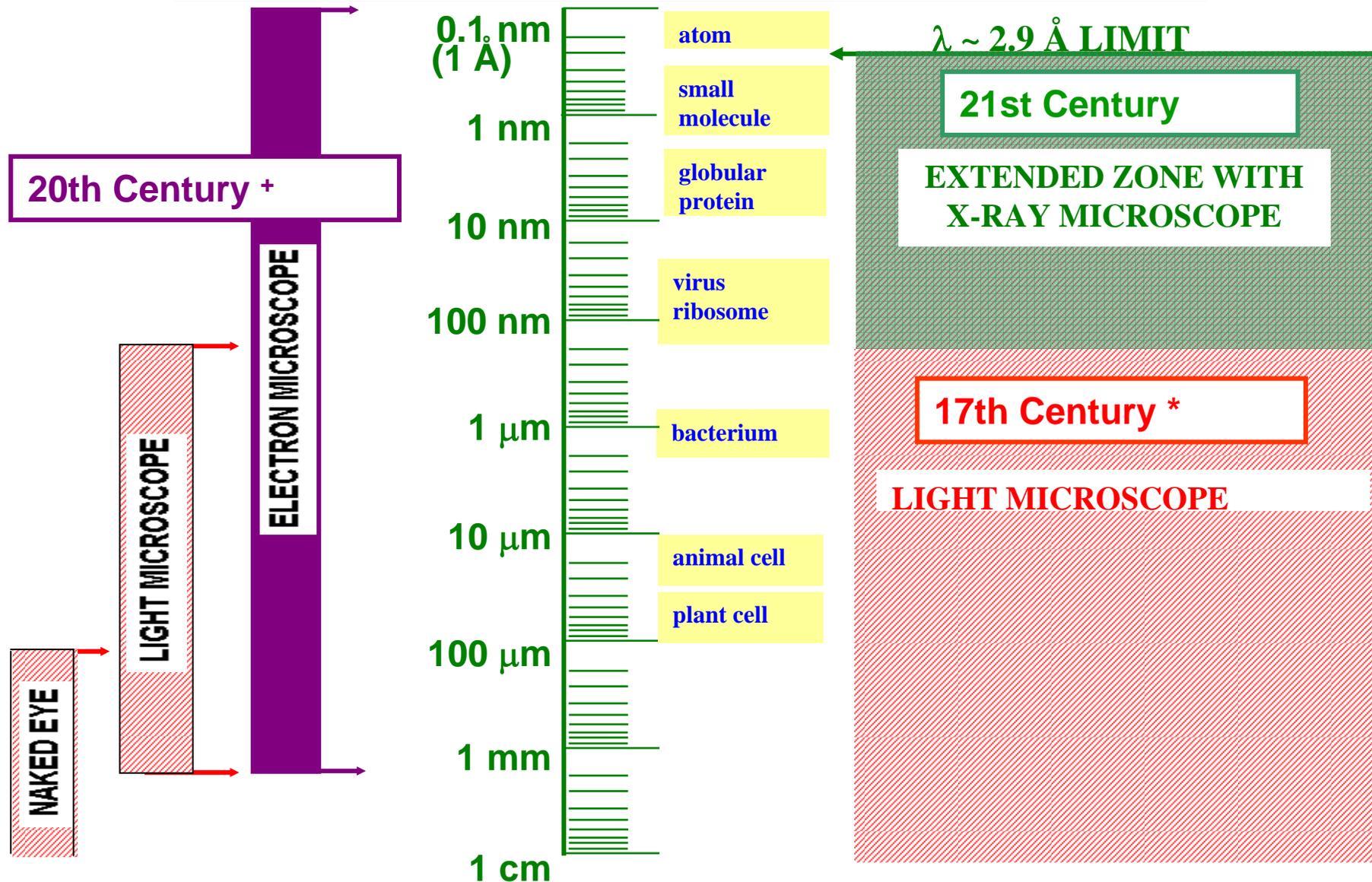


FIG. 2. Stability map in the  $\eta$ - $\rho_0$  plane for relativistic and charge-displacement self-channeling of Gaussian beams in initially homogeneous plasmas. The dimensionless coordinates  $\rho_0$  and  $\eta$  are defined by Eq. 1. Stable and unstable regions in plane  $(\eta, \rho_0)$  and the locus  $\rho_{e,0}(\eta)$  for the lowest eigenmodes are shown. Point  $A_{UNSTABLE}$  (21.1, 20.6) corresponds to the input pulse shown in Fig. 3A. Points  $A_{STABLE}$  and  $B_{STABLE}$  and the trajectory connecting them are described in the text.  $B_{STABLE}$  corresponds to the intensity distribution given in Fig. 3C. (Inset) The exponential longitudinal ( $z$ ) electron density profile [ $N_{e,0}(z) = N_{e,0}(0)\exp(\alpha z)$ ,  $N_{e,0}(z = 3.95 L_R) = 20 N_{e,0}(0) = 1/6 N_{cr}$ ] between points  $A_{STABLE}$  and  $B_{STABLE}$ , and  $L_R$  denotes the Rayleigh range.

# Brightness Comparison



# FOUR CENTURIES OF SPATIAL RESOLUTION LIMITS



• X-RAY IMAGING EXTENDS SPATIAL RESOLUTION OF LIVING MATTER BY ~ 10<sup>3</sup>-FOLD.

\* Kepler, Hooke, Leeuwenhoek

+ Ruska



**LIVING MATTER**



**DEAD MATTER**