History and Future of Laser Technology "Next Generation of Extreme Optical Tools and Applications" On the occasion of 50th Anniversary of Laser Invention 2010 AAAS Annual Meeting San Diego 2/21/10

Relativistic Optics and High Field Science

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International Committee for Ultra Intense Lasers

International promotion of highest intensity lasers and its applications

Inaugurated in 2004, Oxford under IUPAP (above: initial Committee members) Chair: T. Tajima, Co-chair: C. Barty, W. Sandner www.ICUIL.org

Can the society continue to support ever escalating accelerators?



Accelerator = crown-jewel of 20th C science



LHC at CERN



hadron therapy accelerator and gantry

supermagnets quench



Terminated Texas tunnel. The SSC was abandoned after about 25% of the tunnel for the 87-kilometercircumference large collider ring had been bored.



Demise of SSC (Super collider)





Terminated Texas tunnel. The SSC was abandoned after about 25% of the tunnel for the 87-kilometercircumference large collider ring had been bored. By largest machine to probe smallest of structure of matter

size	10^2 km
energy	20TeV
cost	\$10B

US:

Texas site decided (1989)

US Government decided to terminate its work: 1993

Tajima: 'Tamura Symposium' on <u>the Future of Accelerator</u> <u>Physics</u> @ UT Austin (1995)





Dream Beams Symposium

MPQ Garching Feb. 26 – 28, 2007

High intensity laser driven beams

What is <u>collective</u> force ? :Secret behind laser accelerator MU How can a Pyramid have been built?





Individual particle dynamics \rightarrow <u>Coherent</u> and <u>collective</u> movement

Collective acceleration (Veksler, 1956; Tajima & Dawson, 1979)

Collective radiation (N² radiation) Collective ionization (N² ionization)

 \rightarrow Laser driven collective accelerating field



No wave breaks and wake peaks at v≈c



Wave **breaks** at v<c

 $y = X_1 \cos^2 \theta \sin \theta$

 $-\pi/2 < \theta < \pi/2$



Thousand-fold Compactification

Laser wakefield: thousand folds gradient (and emittance reduction?)



The late Prof. Abdus Salam





At ICTP Summer School (1981), Prof. Salam summoned me and discussed about laser wakefield acceleration.

Salam: 'Scientists like me began feeling that we had less means to test our theory. However, with your laser acceleration, I am encouraged'. (1981)

He organized the Oxford Workshop on laser wakefield accelerator in 1982.

Effort: many scientists over many years to realize his vision / dream *High field science*: spawned

Laser technology invented (1985)





Chirped pulse amplification (CPA) invented: to overcome the gain medium nonlinearities in spatially expanded amplification to temporal expansion: smaller, shorter pulse, more intense, higher reprate, all simultaneous.

(Professor Gerard Mourou: ELI Coordinator)

→ many table-top TW and PW lasers world-wide first Chair, ICUIL (International Committee for Ultra Intense Lasers) toward EW laser (*Extreme Light Infrastructure* ELI)

→First LWFA experiments
(Nakajima et al 1994; Modena et al 1995)
→now drives <u>High Field Science</u>

GeV electrons from a centimeter accelerator (a slide given by S. Karsch)





Leemans et al., Nature Physics, september 2006



(Emphasis by S. Karsch)

Key issues of future colliders



(T. Raubenheimer, SLAC, 2008)

Beam Acceleration

- * Largest cost driver for a linear collider is the acceleration
 - − ILC geometric gradient is ~20 MV/m \rightarrow 50km for 1 TeV
- * Size of facility is costly \rightarrow higher acceleration gradients
 - High gradient acceleration requires high peak power and structures that can sustain high fields
 - · Beams and lasers can be generated with high peak power
 - Dielectrics and plasmas can withstand high fields
- Many paths towards high gradient acceleration
 - High gradient microwave acceleration ~ ~100 MV/m
 - Acceleration with laser driven structures
 - Acceleration with beam driven structures
 - Acceleration with laser driven plasmas
 - Acceleration with beam driven plasmas
- SLAC

13th AAC Workshop July 27 - August 2, 2008 ~10 GV/m

Page 11

~1 GV/m



Challenge Posed by DG Suzuki



Frontier science driven by advanced accelerator





compact, ultrastrong a

Can we meet the challenge?

atto-, zeptosecond

A. Suzuki @KEK(2008)



Additional way preparing for the future in Fundamental Physics



- Collider paradigm ('high momentum' approach) quantum mechanics $\Delta E \Delta t \sim \hbar \rightarrow \mathscr{L} \sim E^2$
- Non-collider approaches ('high field' approach) <u>relativity: the higher the energy, the pronounced the effect</u> horizon ~ 1/ *a* (extradimensions?) a = g (Einstein's Equivalence Principle)? Unruh(a)-Hawking(g) radiation? special theory (no preferred frame?; $c(\varepsilon)$?) extreme field physics (merger of research on special and general theories of relativity; Can E also warp space; $c(|E|^2)$) what is vacuum? (QED, QCD(axion), dark energy,...)

(Gies, Marlund, Di Piazza, Dunne, Schuetshold, Heinzl, Reiss, DeKieviert, Rafelski, Zayakin, Smilga, Cohen, Thirolf, Weinfurter, Labun,... discussed)

Quantum Gravity:

"Why is the sky blue?"



(for extreme high energy gamma rays)

 Amelino-Camelia et al., Nature (1998) high energy γ has dispersion:

 $\omega = kc + (extra mass-like term?), i.e. c(\varepsilon)$

- May be regarded as scattering off quantum fluctuations of vacuum (gravitational origin).
- Other proposals, such as H. Sato (1972); Coleman-Glashow(1997),

breakdown of Lorentz invariance?

(cosmic γ rays cease to exist beyond certain energy) May be testable in PeV energy regimes.

Special theory of relativity OK?

doi:10.1038/nature08574

LETTERS

nature

A limit on the variation of the speed of light arising from quantum gravity effects

A list of authors and their affiliations appears at the end of the paper

A cornerstone of Einstein's special relativity is Lorentz invariance the postulate that all observers measure exactly the same speed of light in vacuum, independent of photon-energy. While special relativity assumes that there is no fundamental length-scale associated with such invariance, there is a fundamental scale (the Planck scale, $l_{Planck} \approx 1.62 \times 10^{-33}$ cm or $E_{Planck} = M_{Planck}c^2 \approx 1.22 \times 10^{19}$ GeV), at which quantum effects are expected to strongly affect the nature of space-time. There is great interest in the (not yet validated) idea that Lorentz invariance might break near the Planck scale. A key test of such violation of Lorentz invariance is a possible variation of photon speed with energy¹⁻⁷. Even a tiny variation in photon speed, when accumulated over cosmological light-travel times, may be revealed by observing sharp features in γ -ray burst (GRB) lightcurves². Here we report the detection of emission up to ~31 GeV from the distant and short GRB 090510. We find no evidence for scale (when $E_{\rm ph}$ becomes comparable to $E_{\rm Planck} = M_{\rm Planck}c^2$). For $E_{\rm ph} \ll E_{\rm Planck}$ the leading term in a Taylor series expansion of the classical dispersion relation is $|v_{\rm ph}/c-1| \approx (E_{\rm ph}/M_{\rm QG,n}c^2)^n$, where $M_{\rm QG,n}$ is the quantum gravity mass for order *n* and n = 1 or 2 is usually assumed. The linear case (n = 1) gives a difference $\Delta t = \pm (\Delta E/M_{\rm QG,1}c^2)D/c$ in the arrival time of photons emitted together at a distance *D* from us, and differing by $\Delta E = E_{\rm high} - E_{\rm low}$. At cosmological distances this simple expression is somewhat modified (see Supplementary Information section 4).

Because of their short duration (typically with short substructure consisting of pulses or narrow spikes) and cosmological distances, GRBs are well-suited for constraining $LIV^{2,11,12}$. Individual spikes in long¹³ (of duration >2 s) GRB light-curves (10–1,000 keV) usually show¹⁴ intrinsic lags: the peak of a spike occurs earlier at higher photon-energies. However, there are either no lags or very short lags

(Abdo et al, 2009)



Figure 1 | Light curves of GRB 090510 at different energies. a, Energy

0

0.5

Time since GBM trigger (10 May 2009, 00:22:59.97 UT) (s)

-0.5

1.5

lowest to highest energies. f also overlays energy versus arrival time for each

Meeting Suzuki's Challenge toward PeV $\Delta E \approx 2m_0 c^2 a_0^2 \gamma_{nh}^2 = 2m_0 c^2 a_0^2 \left(\frac{n_{cr}}{n_o}\right), \text{ (when 1D theory applies)}$



	70	(n_e)			
			case I	case II	case III
<i>a</i> ₀			10	3.2	1
energy gain		GeV	1000	1000	1000
plasma density		cm ⁻³	5.7x10 ¹⁶	5.7×10^{15}	5.7×10^{14}
acceleration length	l	m	2.9	29	290
spot radius		μm	32	100	320
peak power		PW	2.2	2.2	2.2
pulse duration		ps	0.23	0.74	2.3
laser pulse energy		kJ	0.5	1.6	5

 $L_{d} = \frac{2}{-}\lambda_{p}a_{0}^{2}\left(\frac{n_{cr}}{r}\right), \qquad L_{p} = \frac{1}{3\pi}\lambda_{p}a_{0}\left(\frac{n_{cr}}{r}\right),$

Even 1PeV electrons (and γ s) are possible, albeit with lesser amount

 \rightarrow exploration of new physics such as the <u>reach of relativity</u> and quantum gravity (correlating with primordial gamma-ray burst [GRB] observation)?

(laser energy of 10MJ@plasma density of $10^{16}/cc$; maybe reduced with index 5/4)





Can we see manifestation of quantum gravity, Lorentz variance in high energy γ? How PeV electrons accelerated?

The Crab Pulsar, a city-sized, magnetized neutron star spinning 30 times a second, lies at the center of this composite image of the inner region of the well-known Crab Nebula. The spectacular picture combines optical data (red) from the Hubble Space Telescope and x-ray images (blue) from the Chandra Observatory, also used in the popular Crab Pulsar movies. Like a cosmic dynamo the pulsar powers the x-ray and optical emission from the nebula, accelerating charged particles and producing the eerie, glowing x-ray jets. Ring-like structures are x-ray emitting regions where the high energy particles slam into the nebular material.

EM Pulse Intensification and Shortening by the Flying Mirror toward the Schwinger field



Hawking radiation





What is 'vacuum'? Does 'something' emerge from 'nothing'? 「空」=「色」? 「混沌」⇔「秩序」? vacuum = 'matter' ? chaos ⇔ information ?

Explore relativity with strong fields (Unruh radiation)

 $A \neq 10^{17} [W/cm^2] \Rightarrow E \approx 10^{12} [V/m]$

(Chen, Tajima 190

 $\Rightarrow k_B T = 0.06 eV \Rightarrow \sim 10 eV$ (blue shift in lab. frame)



Nonlinear Optics in vacuum

Nonlinear optics

in a vacuum

aotonics



LIGHT BULLETS Airy-Bessel wavepackets

ACTIVE PLASMONICS

QUANTUM CASCADE LASERS Efficiency boost What is vacuum? Can vacuum be nonlinear? Is *c* constant? What contribute to nonlinear vacuum?



Detection of (light) fields-particles missed by Collider: exploring new fields such as MU axion.....



A.Chou et al., PRL (2008) observed no signal so far (Note:claim of axion by PVLAS was withdrawn)

High Field Science



(K.Homma)



x[cm]

B. King et al., Nature Photon. 4, 92(2010)



Conclusions

▷ Why strong-field physics ...?

- "...exploring some issues of fundamental physics that have eluded man's probing so far"
 (TAUNACOL)
- QFT: high energy (momentum) vs. high amplitude
- "Fundamental-Physics" discovery potential:
 - ALPs: hypothetical NG bosons (axion, majoron, familon, etc.)
 - MCPs: minicharged particles
 - paraphotons
 - sub-millimeter forces
 -
- high physics/costs ratio

(H. Gies discussed at Extreme Light Infrastructure (ELI) Meeting, 2008)

World Year of Physics 2005

Einstein in the 21st Century

20th Century physics began with Einstein, including theory for laser, 21st Century laser may test and even challenge Einstein.

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Timed to coincide with the 2005 Centennial Celebration of Albert Einstein's *Miraculous Year*, the World Year of Physics 2005 will bring the excitement of physics to the public and inspire a new generation of scientists. Visit www.physics2005.org to find out how you can get involved.

www.physics2005.org

Relativity Helps Acceleration





Strong fields: rectifies laser to longitudinal fields

In <u>relativistic</u> regime, photon x electrons and even protons couple stronger.

> (Tajima, 1999 @LLNL; Esirkepov et al., PRL,2004)

Beyond laser intensity 10²⁴W/cm² ions move relativistically like e⁻

Relativistic and monoenergetic ion beam may constitute compact colliders of ions

 \rightarrow QCD vacuum exploration





some examples:

- Compact cancer hadron therapy devices (JAEA, LIBRA, SAPHIR, Dresden collaborations; will be discussed more)
- Intraoperative Radiation Therapy (IORT): INFN + CEA (Saclay)
- Untrafast radiolysis (LOA etc.)
- Injector for ultrabright X-ray sources (for medicine etc.) (LBL, MPQ etc.)



Conclusions



- Collective acceleration driven by intense laser: leap by many orders (≥ 3), GeV electrons; 10 GeV soon; 100GeV considered; TeV laser collider contemplated; PeV possible?
- <u>High momentum approach</u> vs <u>high amplitude approach</u>: high field science's new paradigm
- Test of Einstein's relativity (special and general theories), nonlinear QED (and QCD), high acceleration (=gravitational) physics, radiation dominant regime, quantum gravity, *nonlinear optics in vacuum*
- Societal applications: already beginning, soon to flourish (e.g., cancer therapy, radiolysis)
- Compact new paradigm of fundamental physics in 21st Century