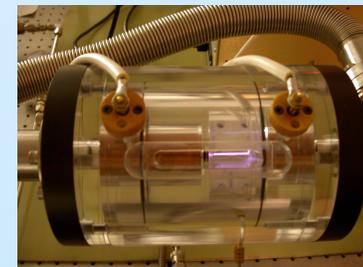
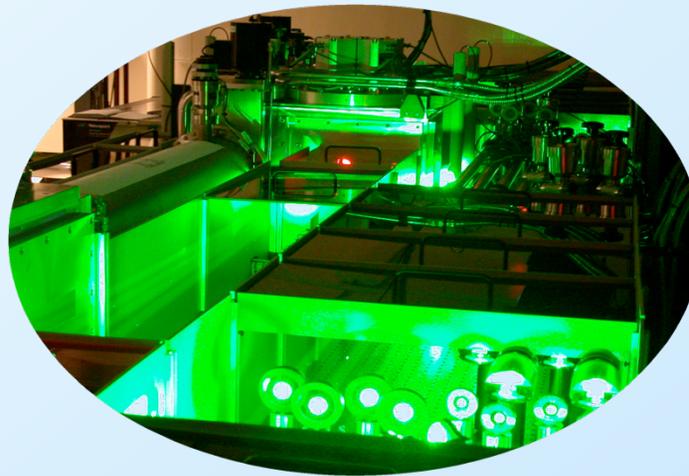


Contrast Enhancement of the LOASIS CPA Laser System and Effects on Electron Beam Performance of LWFA

**Csaba TÓTH¹, Olivier ALBERT^{1,2}, Lorenzo CANOVA², Nicholas MATLIS¹,
Guillaume PLATEAU¹, Anthony GONSALVES¹, Dmitriy PANASENKO¹, and Wim LEEMANS¹**

¹LOASIS Program, Lawrence Berkeley National Laboratory, Berkeley, CA, USA

²Laboratoire d'Optique Appliquée, ENSTA, Ecole Polytechnique, CNRS, Palaiseau, FRANCE



ICUIL'08 – P12
Oct 27-31, 2008 – Shanghai-Tongli, CHINA

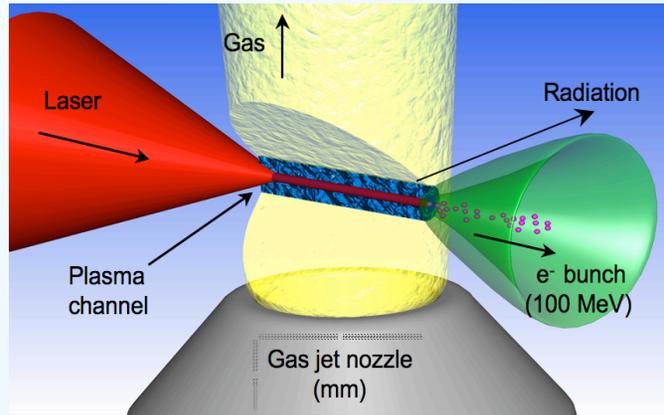
<http://loasis.lbl.gov/>





Applications require stability: driver for hyper-spectral source

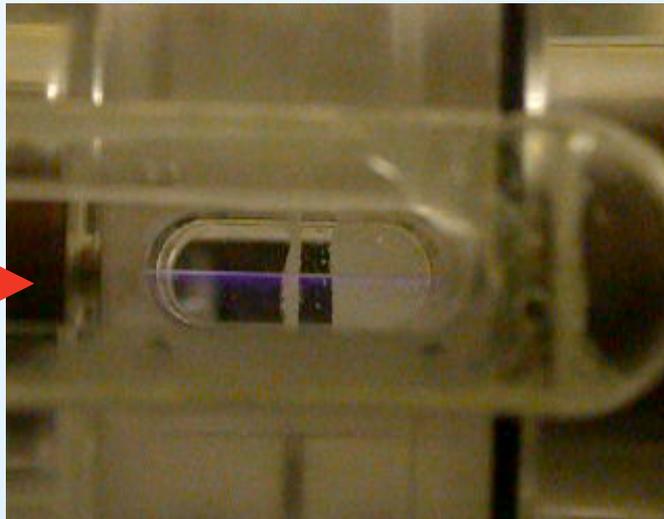
Up to 10 TW
40 fs 10 Hz



150 MeV
e⁻ beam

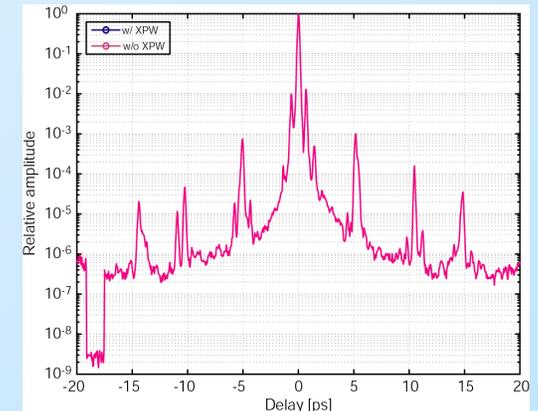
High energy electrons
Compact accelerators

Up to 60 TW
40 fs 10 Hz



~3 cm

1 GeV
e⁻ beam



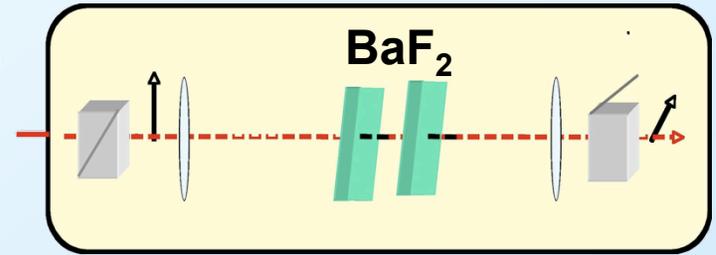
Performance with gas jet affected by laser pre-pulses

Hosokai *et al.*, PRE 2003; PRE2006; Mangles *et al.*, PPCF 2006

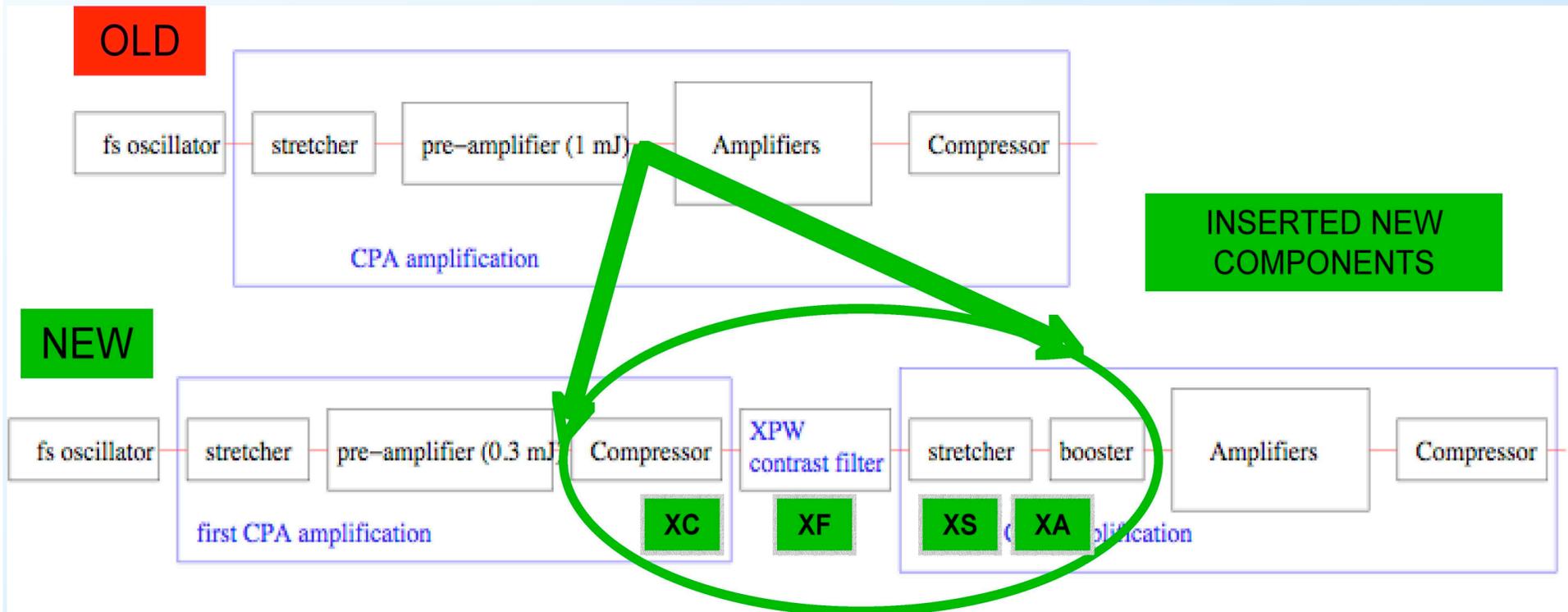


Cross-polarized wave contrast improvement: implementation at LOASIS, LBNL

- Designed by O. Albert (*LOA, France*)
- Non-linear birefringence in BaF_2
- Operates in air, not in vacuum



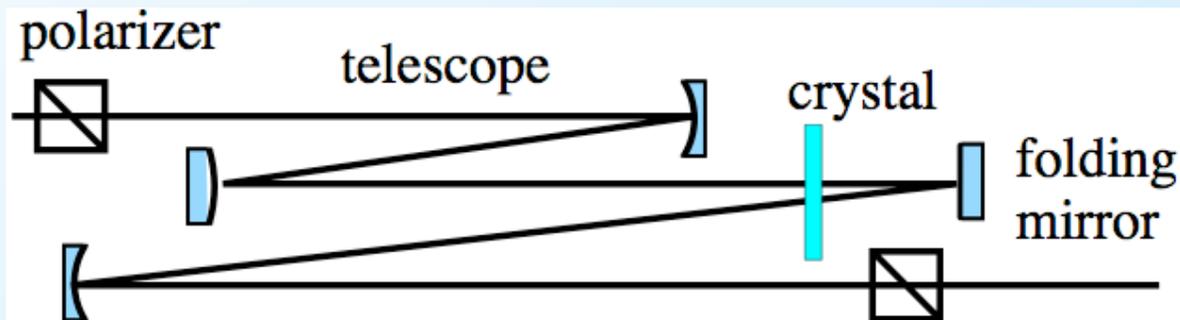
Jullien et al., *Opt. Lett.* 30, 920, 2005



'In-air' high efficiency XPW setup

DESIGN GOALS

- Two crystals XPW scheme for efficiency and reliability [Jullien2006]
- Appropriate input beam matching to maximize XPW efficiency [Albert2006]
- Parameter scaling to avoid SPM in air and SPM in polarizers
- Limit foot-print via folded setup



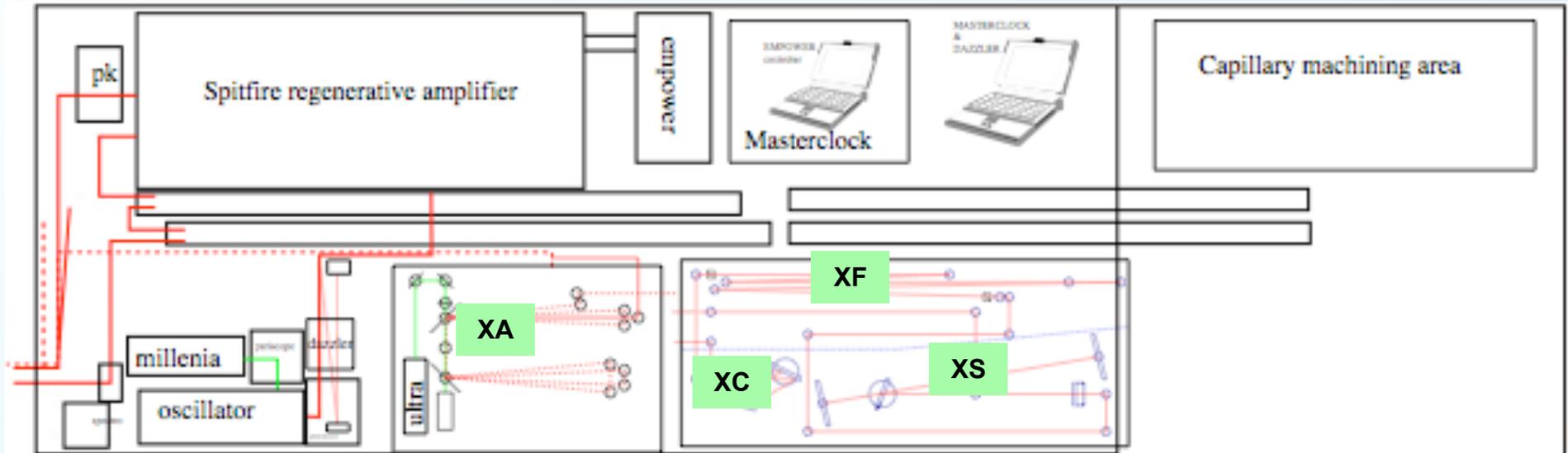
INPUT BEAM PARAMETERS

- 40 fs
- ~ 300 μ J
- Quasi-Gaussian mode
- 6 mm beam diameter

DESIGN PARAMETERS

- Focusing telescope : 4 m focal length
- BaF₂ crystal spacing ~ 60 cm
- Output beam diameter : ~ 2 mm
- 50 μ J output in everyday operation
- 25% max measured efficiency
- Small footprint 1.5 x 0.2 m²

XPW layout: implementation at LOASIS



new beam pass
present beam pass

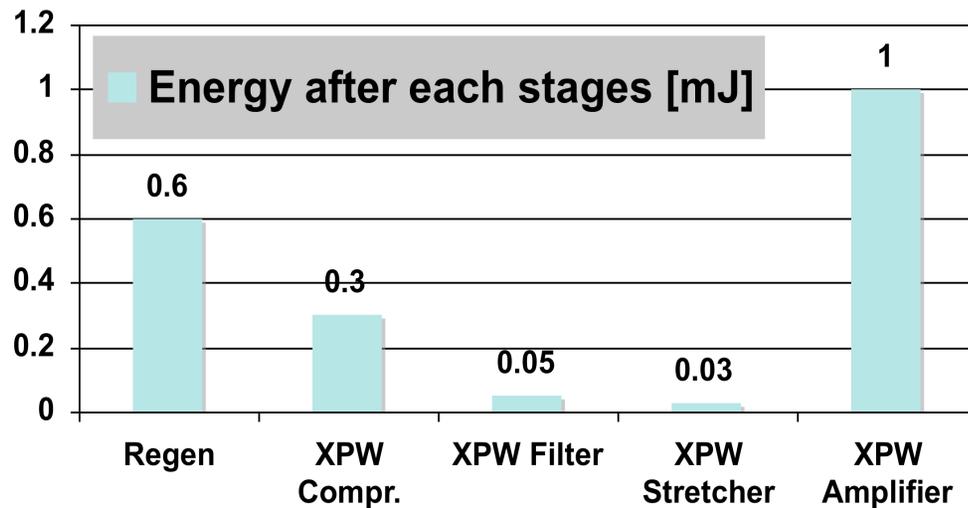
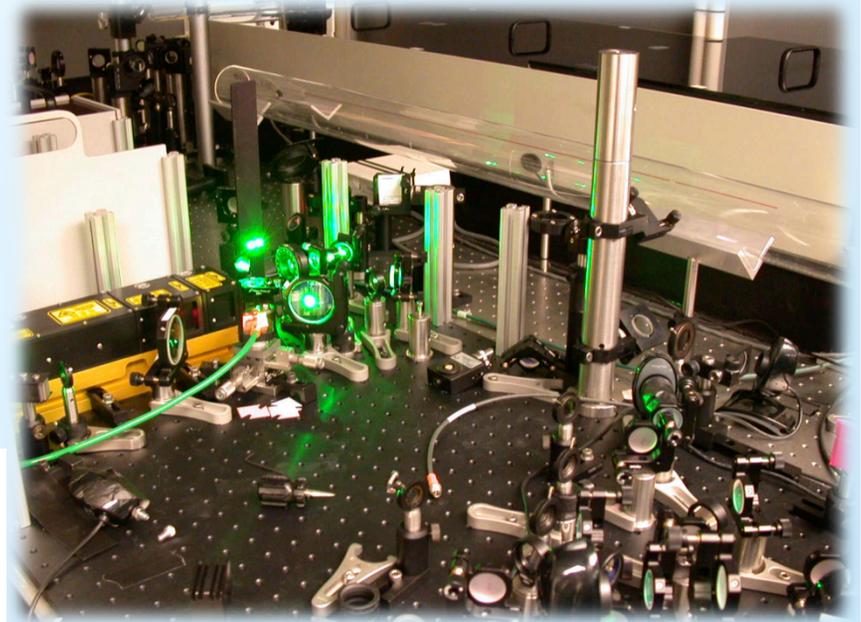
MAIN FEATURES

- 17% efficiency ($300 \mu\text{J} \rightarrow 50 \mu\text{J}$)
- Spectral broadening: 25 nm to 41 nm
- Gaussian beam mode output
- Energy fluctuation $\leq 5\%$
- Optimized XPW design: XPW setup in air, no need to be in vacuum (adapted input beam diameter to avoid SPM in polarizer, good focal length to avoid SPM in air at focus)
- Folded design: small footprint (1.5 m x 0.2 m)

XC - XPW COMPRESSOR
XF - XPW FILTER
XS - XPW STRETCHER
XA - XPW AMPLIFIER

Second CPA & Booster Amplifier

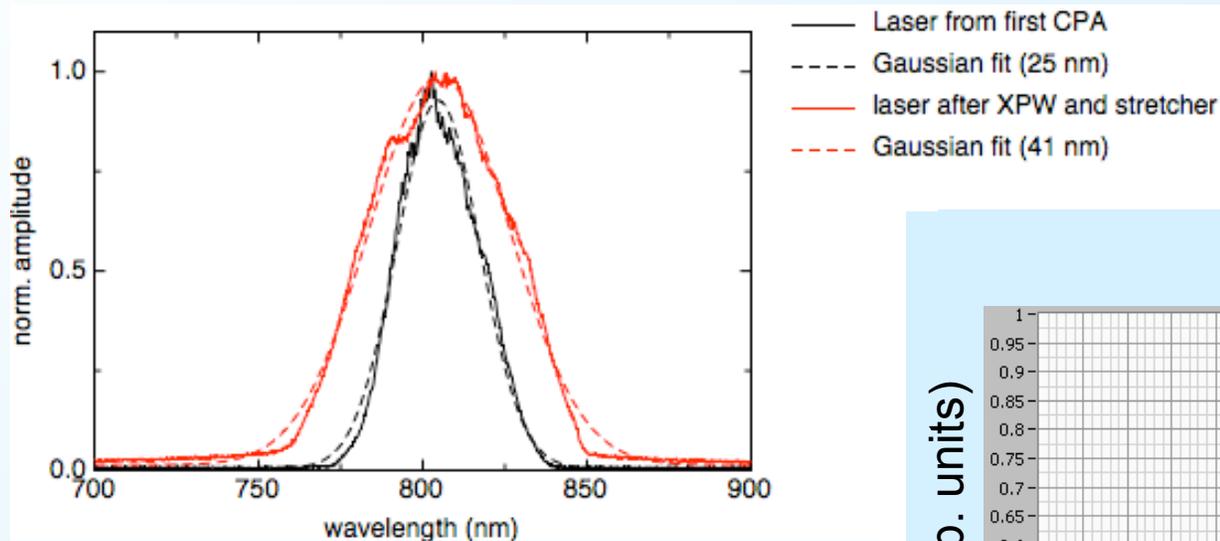
- Contrast preservation requirements:
 - Pump fluence: 1 to 1.4 J /cm²
 - Passes with gain below 4/pass
 - Total 4 passes



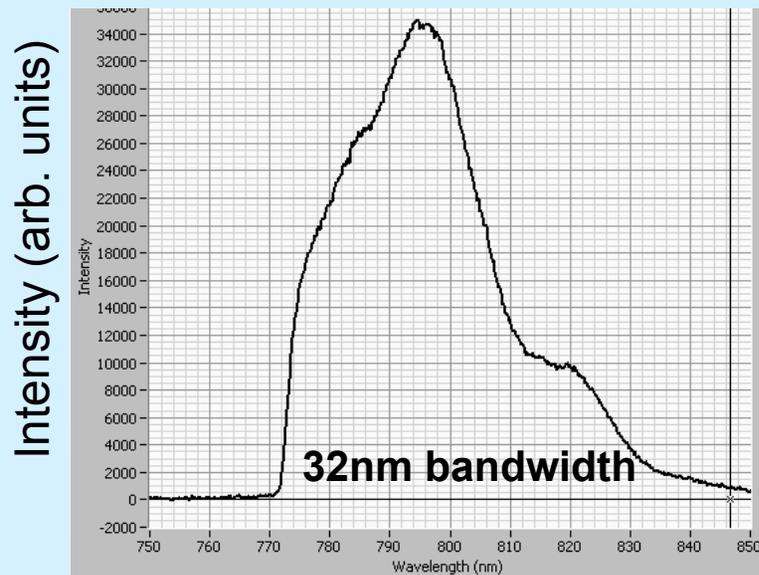
‘Amp0’ with pump laser –
‘Ultra’ from Big Sky Lasers



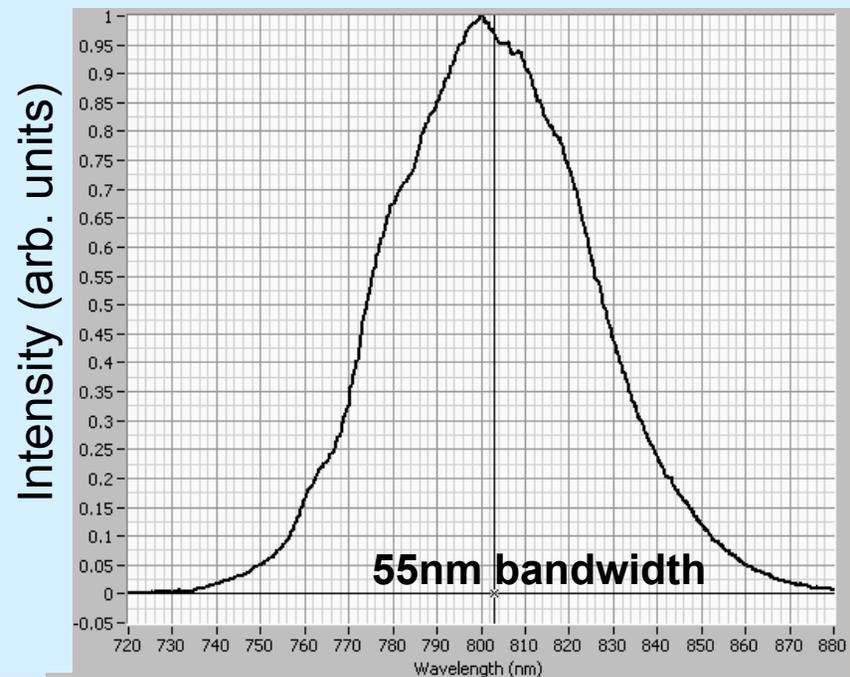
Spectrum broadened after XPW but very sensitive to compressor alignment



Misaligned



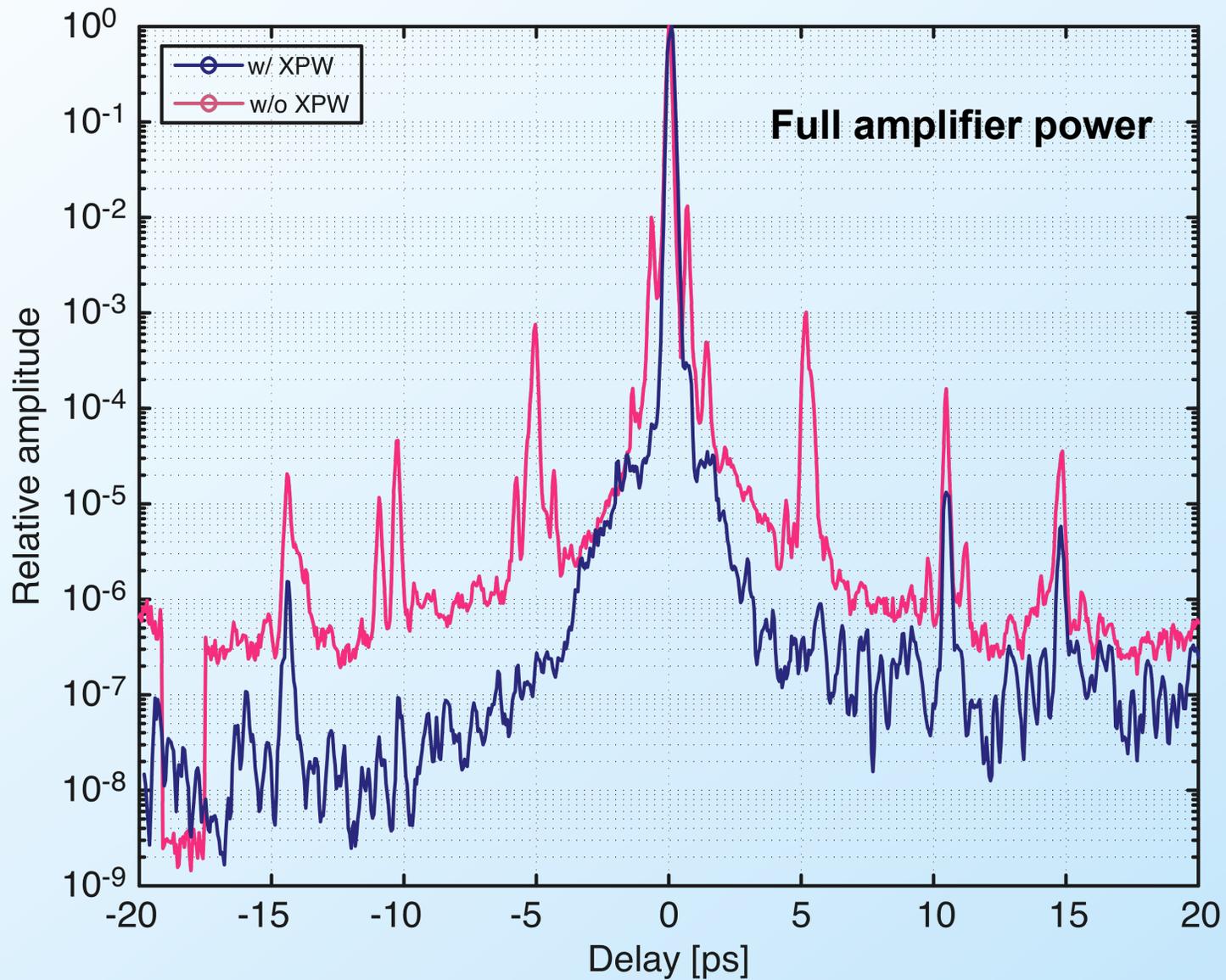
Aligned



Wavelength (nm)



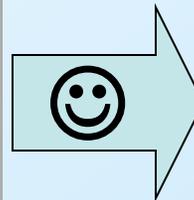
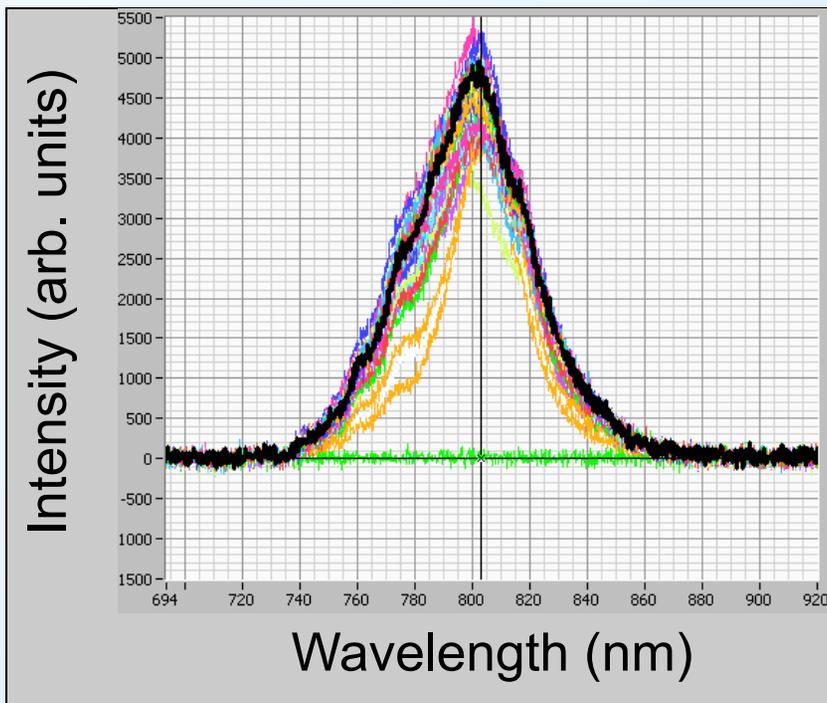
XPW reduces pre-pulse by 3-4 orders



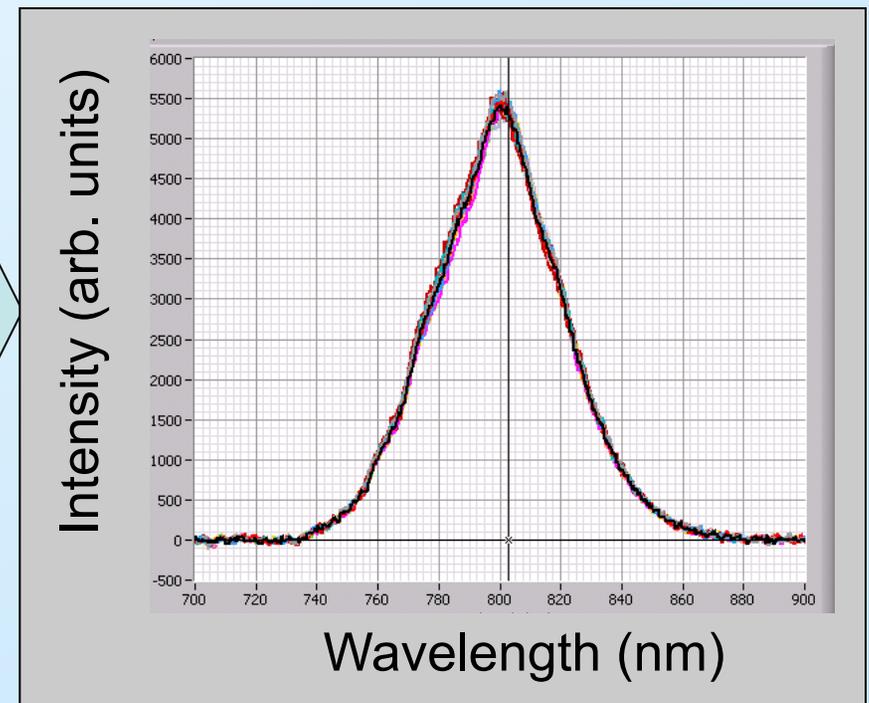
XPW stability

- Spectrum and energy was unstable
 - Turned down input power to make stable (+ some minor tuning of compressor)

20080318



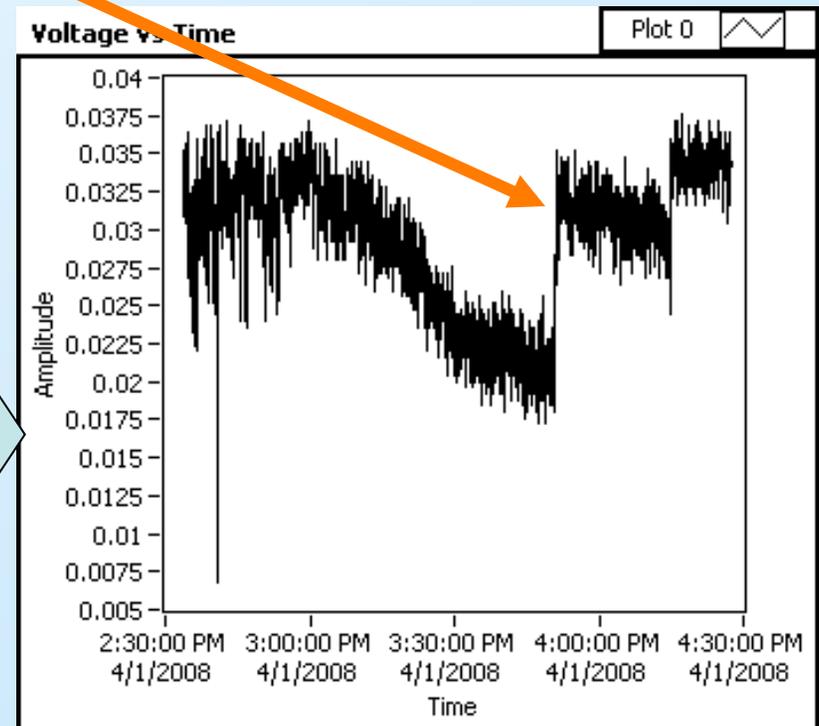
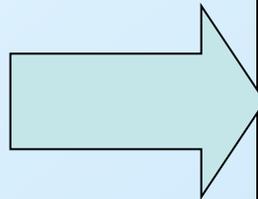
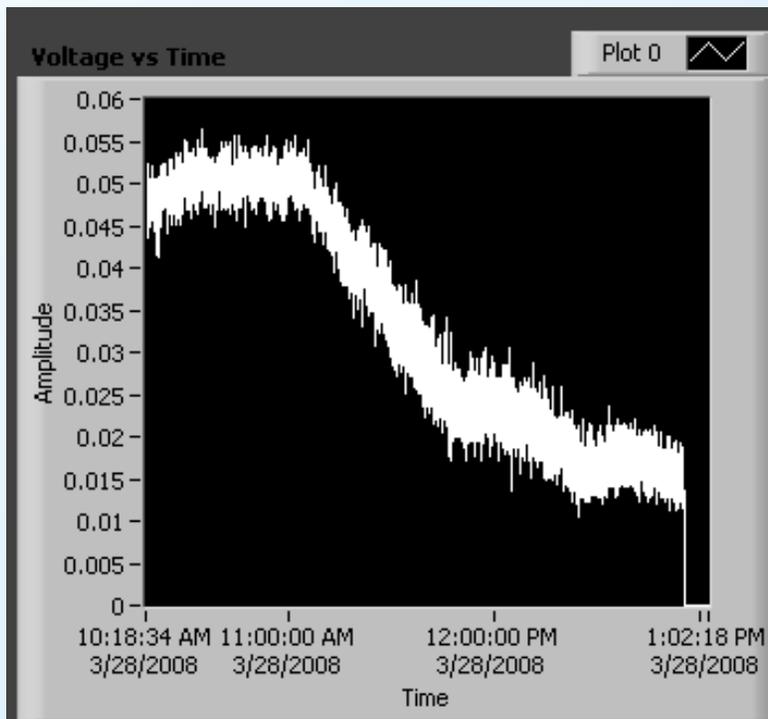
20080404





Booster amplifier stability strongly affected by pointing due to small gain volume

- System has been equipped with motorized controls
- Closed feedback installation in process but drift slow enough for operator to give occasional correction kick

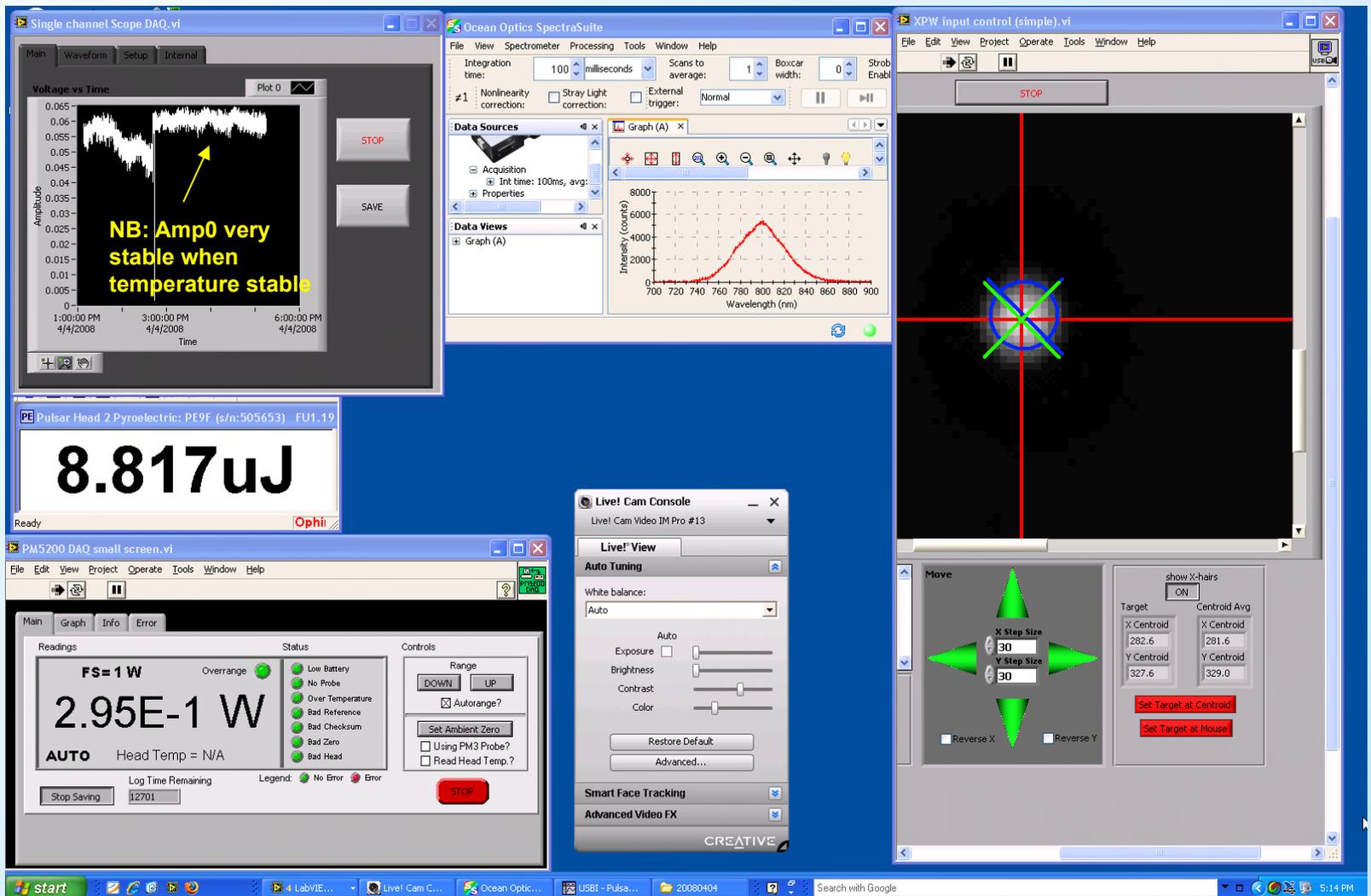


XPW logging and control summary

Amp0 energy

XPW output spectrum

XPW pointing to Amp0



Single channel Scope DAQ.vi
 Voltage vs Time
 NB: Amp0 very stable when temperature stable

Ocean Optics SpectraSuite
 Integration time: 100 milliseconds
 Scans to average: 1
 Boxcar width: 0
 Graph (A) showing Intensity (counts) vs Wavelength (nm)

XPW input control (simple).vi
 STOP button
 Graph (A) showing a target area with a blue circle and red crosshairs

PE Pulsar Head 2 Pyroelectric: PE9F (s/n:505653) FU1.19
8.817uJ

PH5200 DAQ small screen.vi
 FS = 1 W
2.95E-1 W
 Head Temp = N/A

Live! Cam Console
 Live! Cam Video IM Pro #13
 Auto Tuning: White balance: Auto
 Exposure, Brightness, Contrast, Color sliders
 X Step Size: 30, Y Step Size: 30

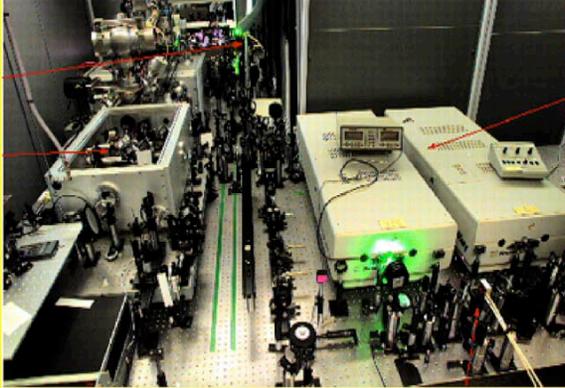
Regen energy

Preamp1 power

Windows taskbar: LabVIEW, Live! Cam C..., Ocean Optic..., USB1 - Pulsa..., 20080404, Search with Google, 5:14 PM

LOASIS multi-terawatt laser systems

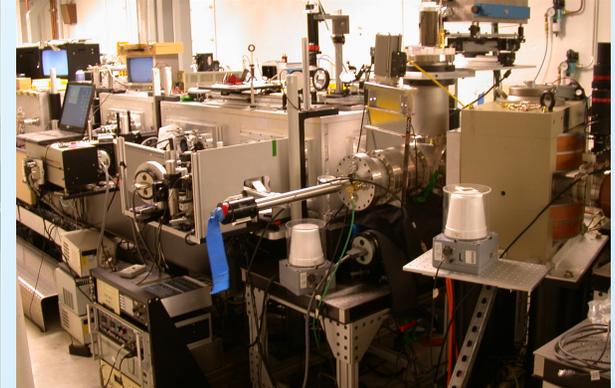
10 TW Ti:sapphire



TREX laser



Shielded target room



Three main amplifiers (Ti:sapphire, 10 Hz):

- “Godzilla”:

0.5-0.6 J in 40-50 fs (10-15 TW) ==> drive beam for **gasjet exps**

- “Chihuahua”:

20-50 mJ in 50 fs

==> ignitor beam } guiding

250-300 mJ in 200-300 ps

==> heater beam

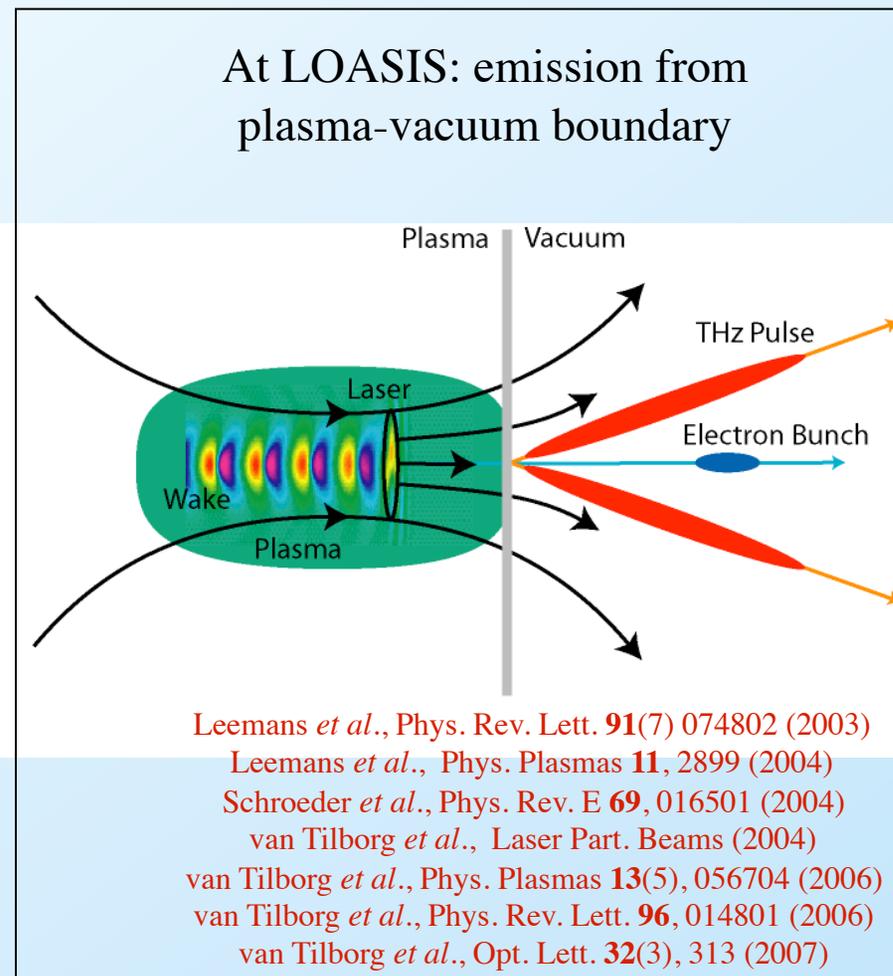
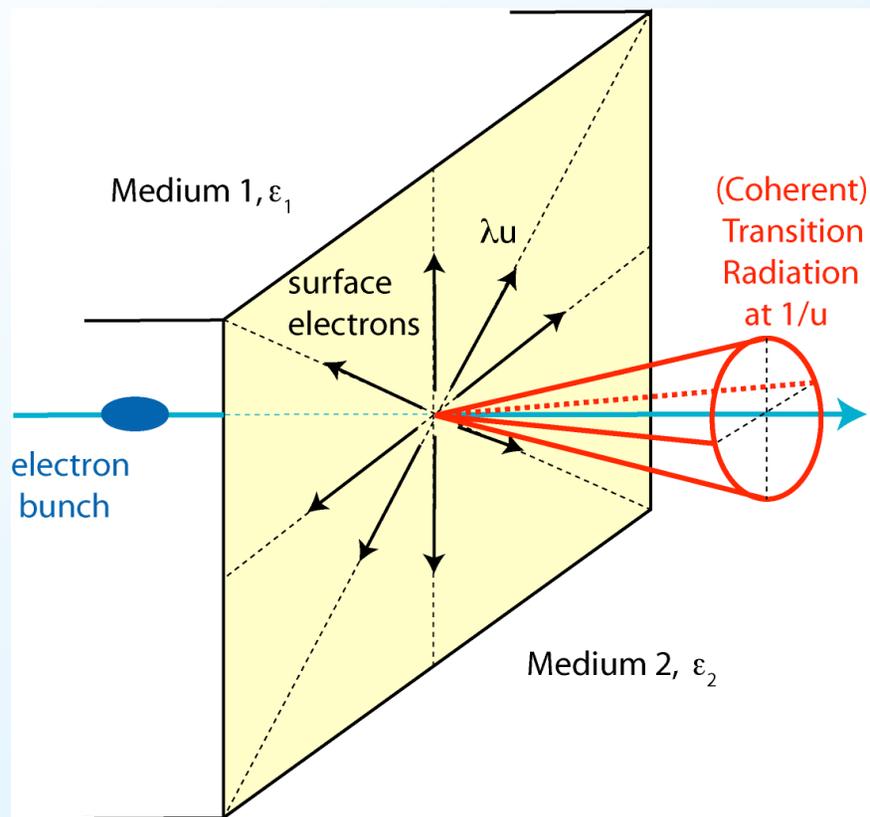
20-80 mJ in 50 fs

==> colliding beam

- “TREX”:

2.7 J in 35-40 fs (60 TW) ==> capillary guiding exps

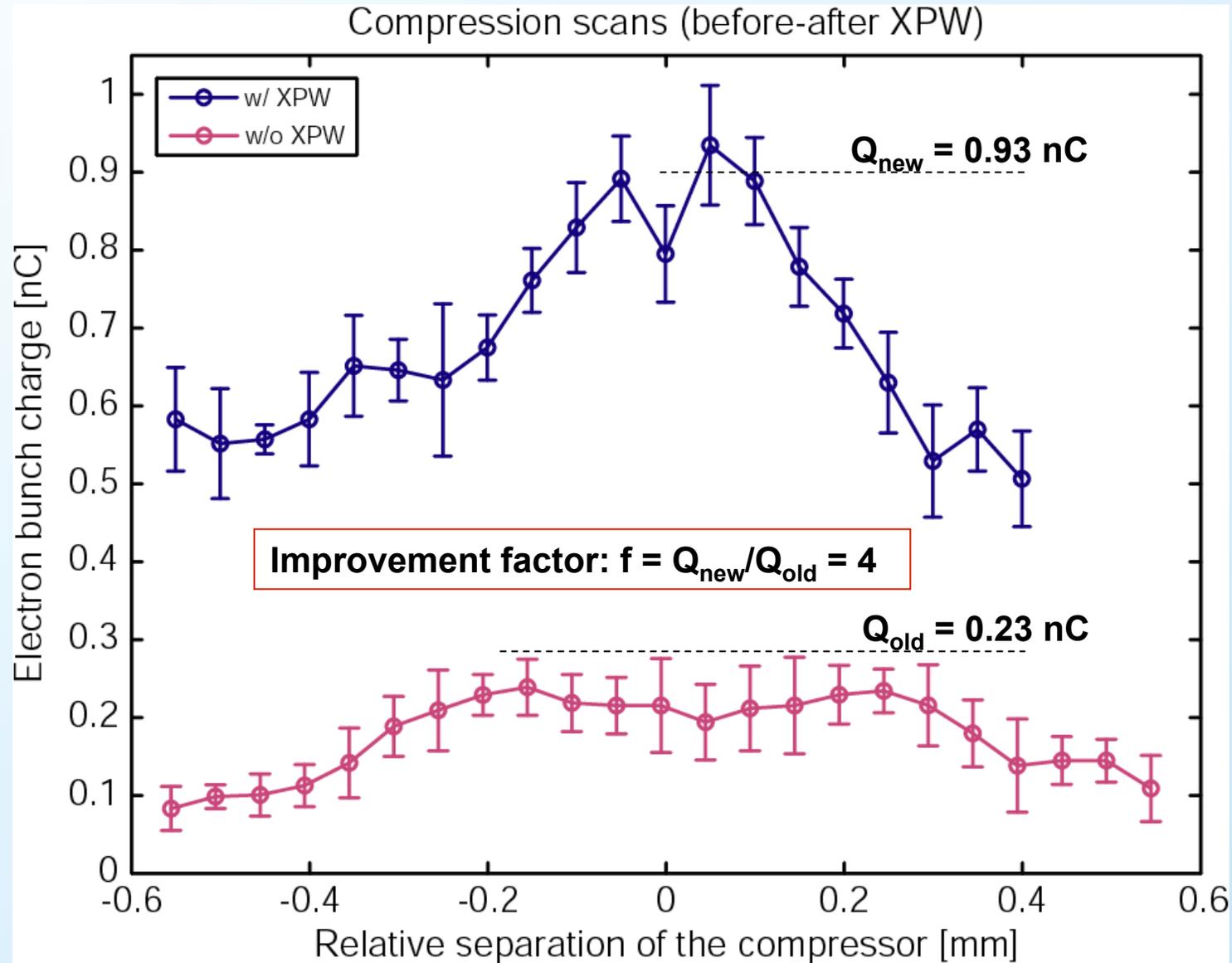
Application: improving laser accelerators



$$\frac{d^2 I_{e^-}}{d\omega d\Omega} = \frac{e^2}{\pi^2 c} \left| \frac{u\sqrt{1+u^2} \sin \theta}{1+u^2 \sin^2 \theta} \right|^2$$

θ = emission angle
 u = electron momentum

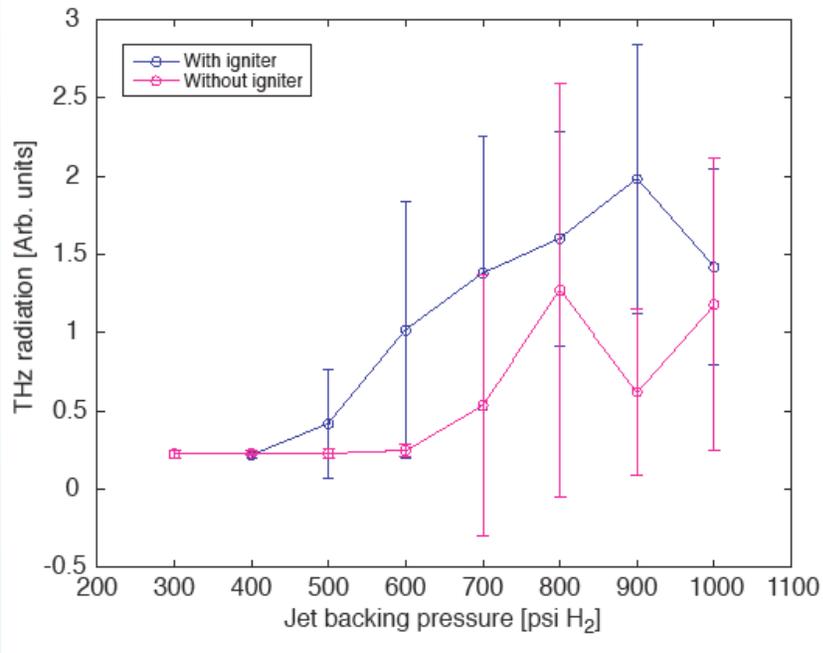
Improved contrast quadruples charge yield



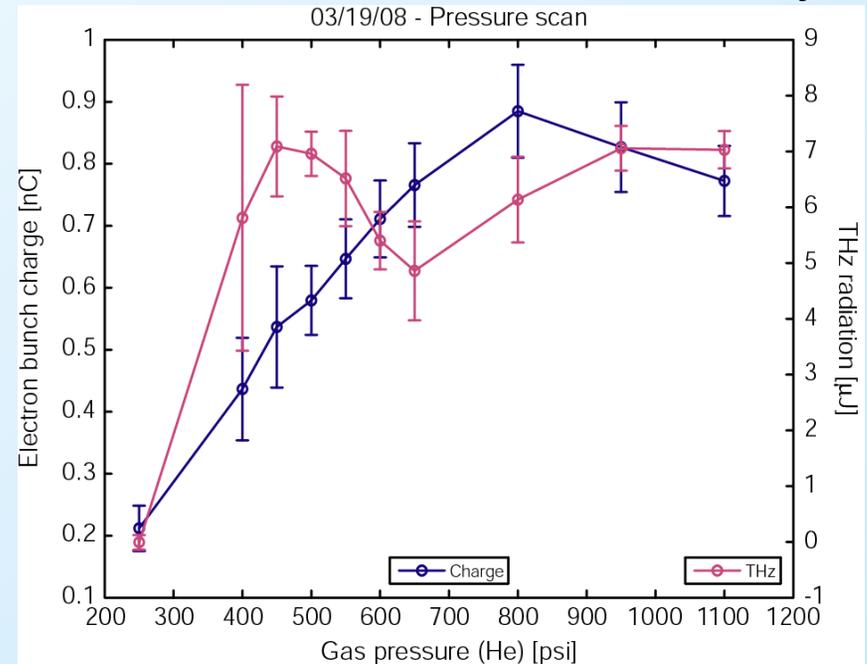


Improved pulse contrast increases THz stability

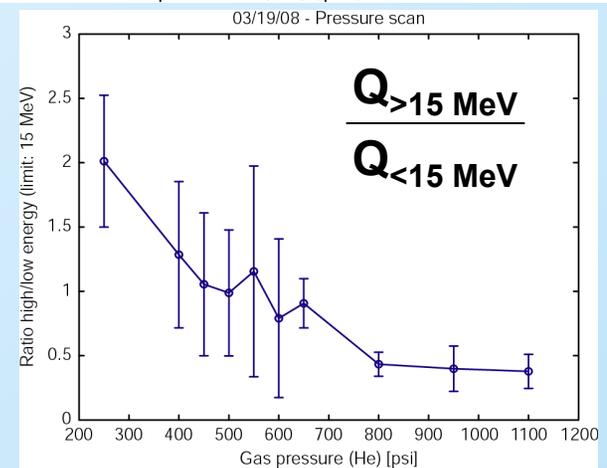
Pre-XPW: 50 – 100% variability



Post-XPW: 10 – 20% variability



Pressure scan shows transition from high-energy (*resonant*) regime to low-energy (*self-modulated*) regime





Conclusion

- Implemented XPW based contrast improvement
 - In-air design
 - Compact system
- Contrast improved by 3-4 orders
- Spectrum broadened from ~25 nm to ~40 nm
- Beam pointing stabilization essential to maintain stable operation of XPW
- Substantial improvement in operation of laser accelerator based on gas jet